

Topic 2: Mechanics

Content:

- mechanics - basic terms and quantities
- velocity and acceleration
- force, moment of force, momentum
- work, power
- Newton's laws
- Kepler's laws
- Newton's gravitational law
- free fall, motion in gravitational field
- mechanical energy

Classical mechanics

ČIN VYSVETLUJE GALILEIHO 57

idelí nijakú rýchlosť, len
stane pri ňom bez
ento jav si vysvetľuje
vačnosti, podľa ktorého
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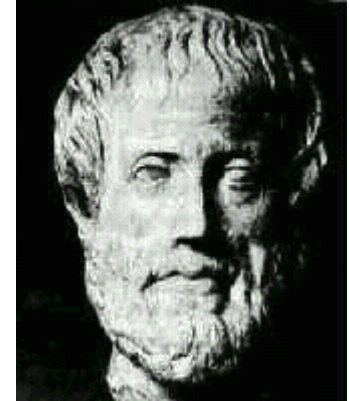
sti som mal

ej relativity a jej
sebe Einstein
muláciu
šak bolo treba
to povedať, že
chápali jej
podobne



The Ancient Greeks

Aristotle (384-322 B.C.) is regarded as the first person to attempt physics, and actually gave physics its name.



On the nature of matter:

Matter was composed of:

Air

Earth

Water

Fire

Every compound was a mixture of these ***elements***

There was **no predictive aspect**

(science should have it – from the todays view)

and no need to verify statements by experiments.

On the Nature of Motion

- **Natural** motion - like a falling body
 - Objects seek their natural place
 - Heavy objects fall fast
 - Light objects fall slow
- **Unnatural** motion - like a cart being pushed
 - The moving body comes to a stand still when the force pushing it along no longer acts
 - The natural state of a body is at rest.

Aristotelian Physics

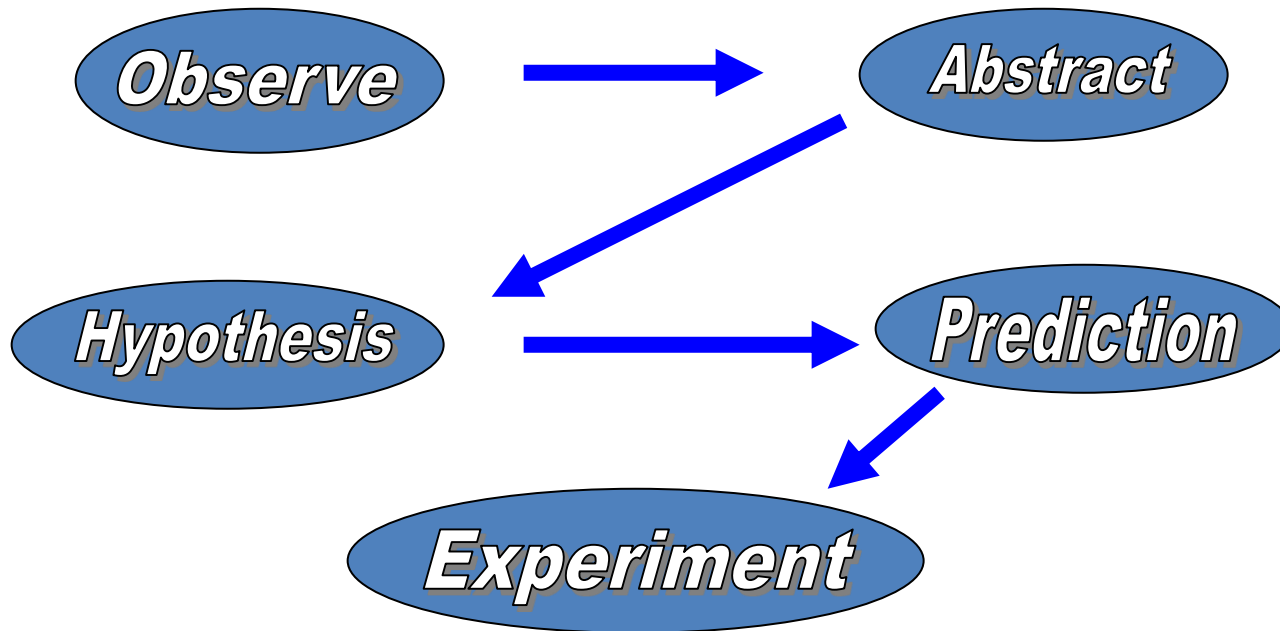
- Aristotelian Physics was based on logic
 - o It provided a framework for understanding nature
 - o It was logically consistent

but it was wrong !!!
(from todays view)

- Aristotelian physics relied on logic - not experiment!

The Renaissance

Galileo Galilei (1564 -1642) was one of the first to use the scientific method of observation and experimentation. He laid the groundwork for modern science.



Classical Mechanics

Mechanics: the study of *motion*

Galileo (1564 -1642) laid the
groundwork for Mechanics

Newton (1642-1727) completed
its development (*~almost~*)

Newton's Laws work fine for



- Large Objects - Ball's, planes, planets, ...
 - Small objects (atoms) → Quantum Mechanics
- Slow Objects - people, cars, planes, ...
 - Fast objects (near the speed of light) → Relativity
- Classical Mechanics - essentially complete at the end of the 19th Century

Later during the summer term we will come also
to specific/general relativity and quantum mechanics.

Classical mechanics

Galileo's vs Newton's understanding of inertia:

Galileo:

- developed the idea of force, as a cause for motion
- determined that the natural state of an object is rest or uniform motion, i.e. objects always have a velocity, sometimes that velocity has a magnitude of zero = rest.
- objects resist change in motion, which is called **inertia**

Newton (further and more detailed understanding):

- change in velocity = acceleration caused by force
- **inertia** = resistance to change in motion – it is proportional to the mass of the object and the acting force.

mechanics - basic terms and quantities

The general study of the relationships between motion, forces, and energy is called **mechanics**.

Motion is the action of changing location or position.

Motion may be divided into three basic types - translational, rotational, and oscillatory.

The study of motion without regard to the forces or energies that may be involved is called **kinematics**. It is the simplest branch of mechanics.

The branch of mechanics that deals with both motion and forces together is called **dynamics** and the study of forces in the absence of changes in motion or energy is called **statics**.

basic terms and quantities

point (point mass) is a mass which doesn't have volume (mass focused into one point),

body (physical object) is an identifiable collection of matter, which may be more or less constrained to move together by translation rotation, or oscillation,

rigid body is an idealization of a solid body in which deformation is neglected,

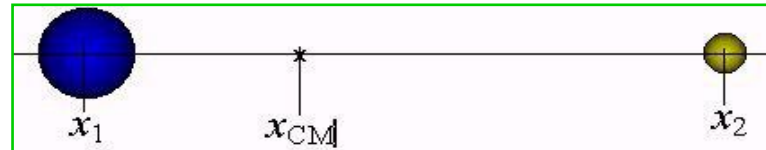
Soft body



Rigid body



barycentre is the center of mass of two or more bodies that are orbiting each other,



trajectory - path that a moving object follows through space as a function of time.

basic terms and quantities

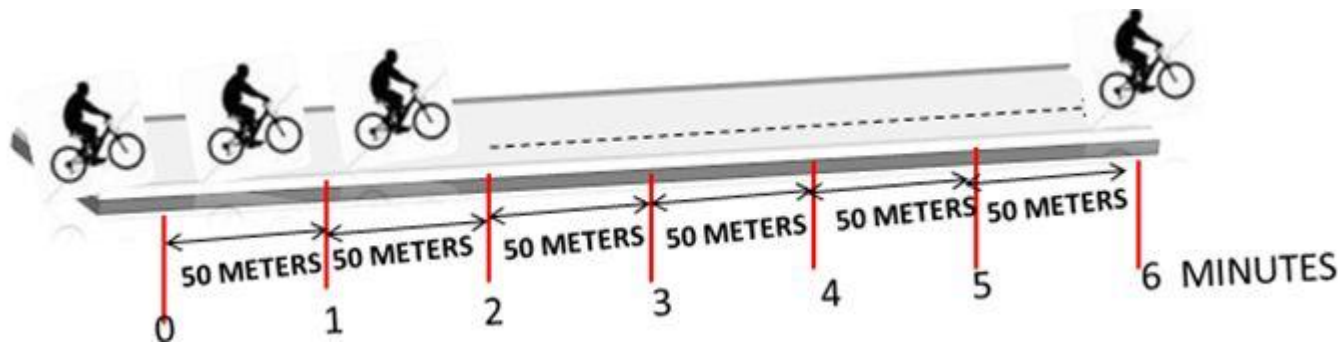
Velocity: The velocity $\vec{V}$ of an object is the rate of change of its position with respect to a frame of reference, and is a function of time.

It is in general a vector quantity.

Its size (scalar value = speed) can be evaluated:

$$|\vec{V}| = v = \frac{\Delta s}{\Delta t} \approx \frac{ds}{dt} = s' \quad [\text{m} \cdot \text{s}^{-1}]$$

where s is path (distance) [m] and t is time [s].



velocity vs speed

velocity vs. speed

velocity is vector, speed is scalar

basic terms and quantities

Acceleration: is the rate of change of velocity (speed) of an object.

It is in general a vector quantity.

Its size can be evaluated:

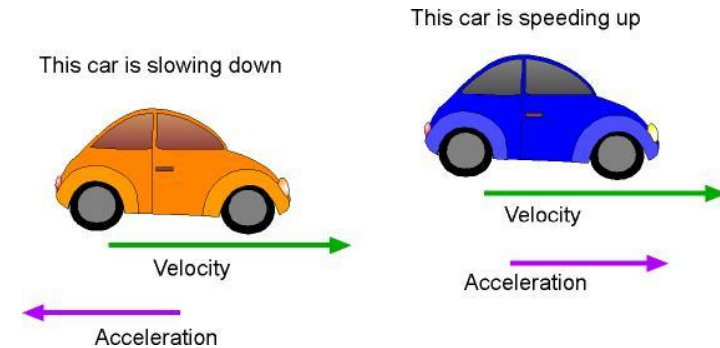
$$|\vec{a}| = \frac{\Delta v}{\Delta t} = \frac{dv}{dt} = \frac{d^2s}{dt^2} = v' = s'' \left[m \cdot s^{-2} \right]$$

v is speed, s is path (distance) [m], and t is time [s].

Important equations:

$$v = v_0 + at \quad s = v_0 t + \frac{1}{2}at^2$$

where v_0 is the so-called starting velocity.



Tangent vs normal component

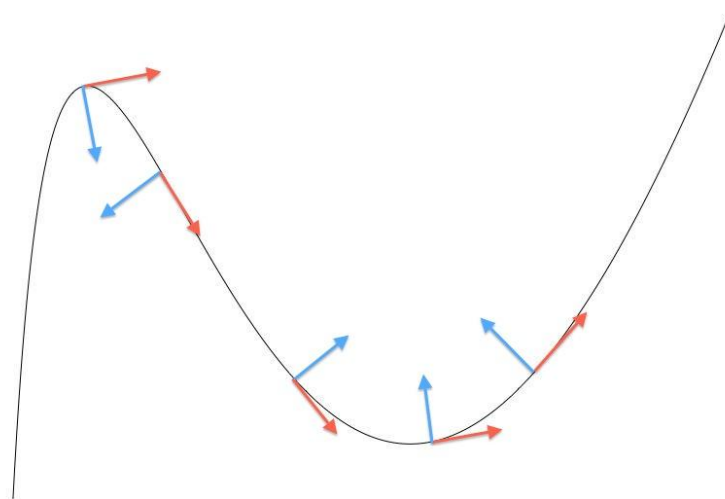
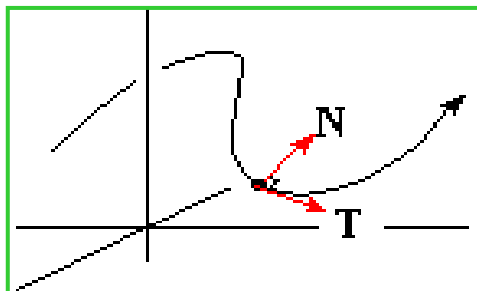
- each vector of velocity and/or acceleration can be split into 2 perpendicular components along a curvilinear path:

- **tangent** component (v_t)

(connected with the change of the vector size)

- **normal** component (v_n)

(connected with the change of vector direction)



basic terms and quantities

Angular velocity: is the rate of change of angular displacement (angle) as a function of time.

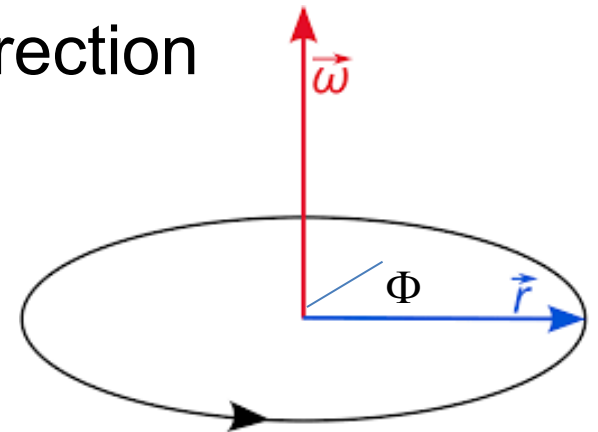
It is defined for rotating bodies (because tangential velocity is a function of radius).

It is a vector quantity (pointing in the direction of rotation axis).

Symbol: $\vec{\omega}$

$$|\vec{\omega}| = \frac{\Delta\phi}{\Delta t} = \frac{d\phi}{dt}$$

Unit: radians per second [$\text{rad}\cdot\text{s}^{-1}$]



Relationship between angular and

tangential velocity: $|\vec{v}_t| = |\vec{\omega}|r$ ($\vec{v}_n = 0$)

And normal acceleration: $|\vec{a}_n| = |\vec{\omega}|^2 r$ ($\vec{a}_t = 0$)

classification of motions

a_t	a_n	a	motion
$a_t = 0$	$a_n = 0$	$a = 0$	uniform straight-line motion; velocity has constant direction and size
$a_t = 0$	$a_n \neq 0$	$a \neq 0$	uniform curvilinear motion; velocity has constant size, but the direction is changed (e.g. motion along a circular line)
$a_t = \text{const.}$	$a_n = 0$	$a \neq 0$	uniformly accelerated, straight-line motion, velocity has the same direction, velocity size is changed
$a_t = \text{const.}$	$a_n \neq 0$	$a \neq 0$	uniformly accelerated, curvilinear motion, velocity direction is changed, velocity size is changed
$a_t \neq \text{const.}$	$a_n = 0$	$a \neq 0$	non-uniformly accelerated, straight-line motion, velocity has the same direction, velocity size is changed
$a_t \neq \text{const.}$	$a_n \neq 0$	$a \neq 0$	non-uniformly accelerated, curvilinear motion, velocity direction is changed, velocity size is changed
		$a = \text{const.}$	motion with constant acceleration (e.g. free fall in Earth gravity field)
a - total acceleration, $a = a_t + a_n$, a_t - tangent component of acceleration, a_n - normal component of acceleration,			

basic terms and quantities

Mass (weight):

Mass is both a *property* of a physical body and a *measure of its resistance* to acceleration (a change in its state of motion) when a force is applied.

- inertial mass measures an object's resistance to being accelerated by a force (represented by the relationship $F = ma$).
- gravitational mass (weight) measures the gravitational force exerted on an object in a known gravitational field.

Comment: Mass of an object should be the same on the Earth and Mars, but the weight will be different.

Unit: [kg]

Its definition has to be changed due to the problems with the international prototype.

Mass (weight) is a scalar quantity.

mass vs. weight

mass vs. weight

mass is independent from gravity field

basic terms and quantities

Force: is any interaction that, when unopposed, will change the motion of an object (Galileo).

It will change its velocity (Newton).

If the mass of the object is constant, this law implies that the acceleration of an object is directly proportional to the force acting on the object (in the direction of the force).

Its size can be evaluated (2. Newton's law):

$$\vec{F} = m\vec{a} \quad [\text{kg} \cdot \text{m} \cdot \text{s}^{-2}] = [\text{N}]$$

where m is mass [kg] and $\mathbf{a}$ is acceleration [$\text{m} \cdot \text{s}^{-2}$].

It is a vector quantity.

Unit [N] is called Newton.

basic terms and quantities

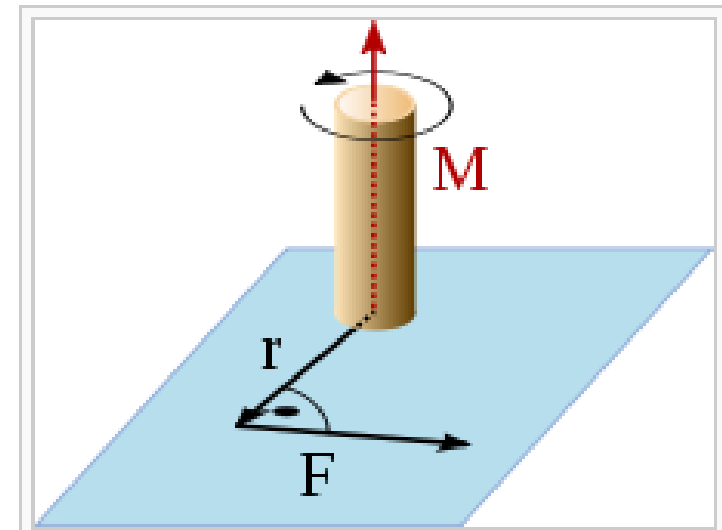
Moment of force (torque): is the product of a force and its distance from an axis, which measures the rotation effect of the force (about that axis).

In general, it is a combination of a physical quantity and a distance.

$$\vec{M} = \vec{F} \times \vec{r} \quad [\text{N} \cdot \text{m}] = [\text{kg} \cdot \text{m}^2 \cdot \text{s}^{-2}]$$

where $\vec{F}$ is force [N] and
 $\vec{r}$ is distance vector [m].

It is in general a vector quantity.

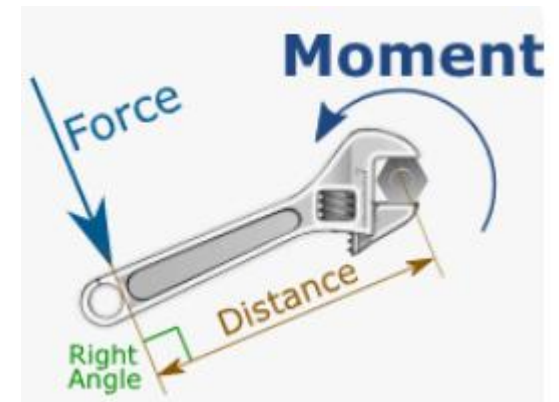
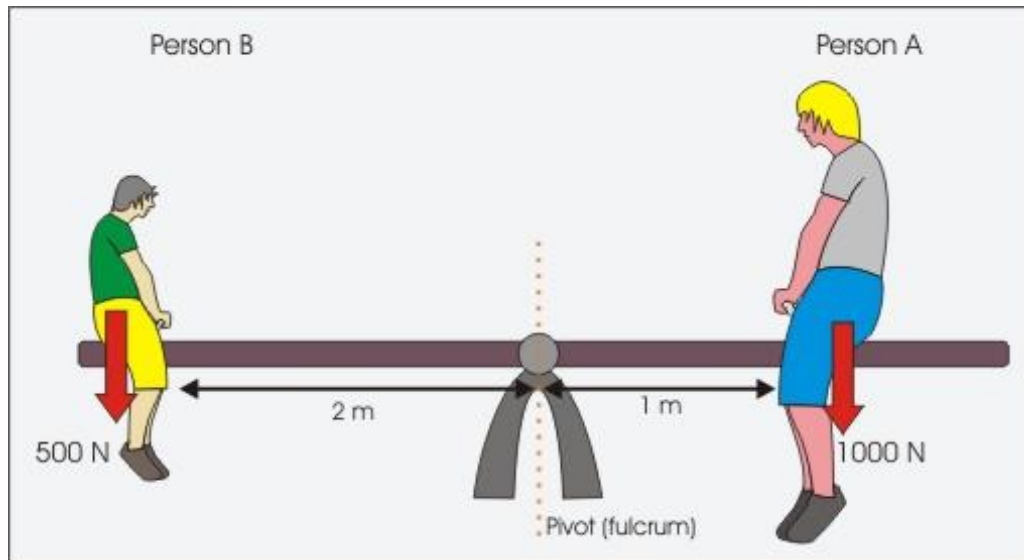
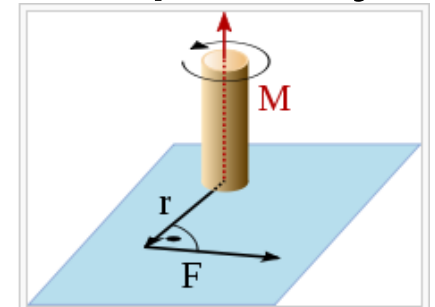


basic terms and quantities

Moment of force (torque): is the product of a force and its distance from an axis, which measures the rotation effect of the force (about that axis).

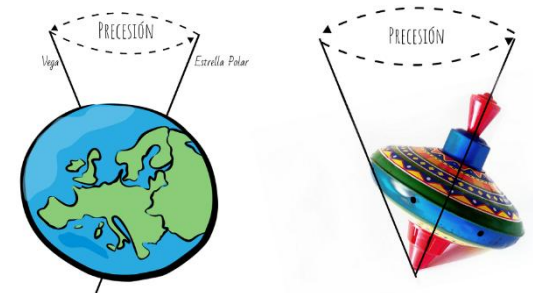
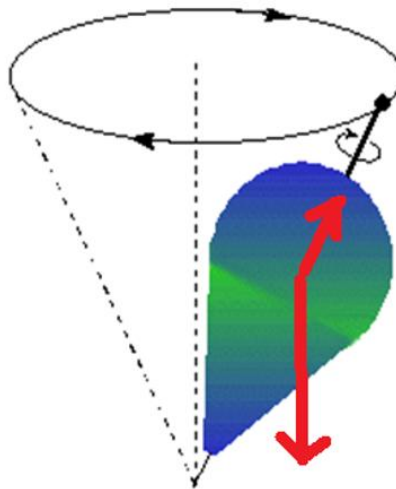
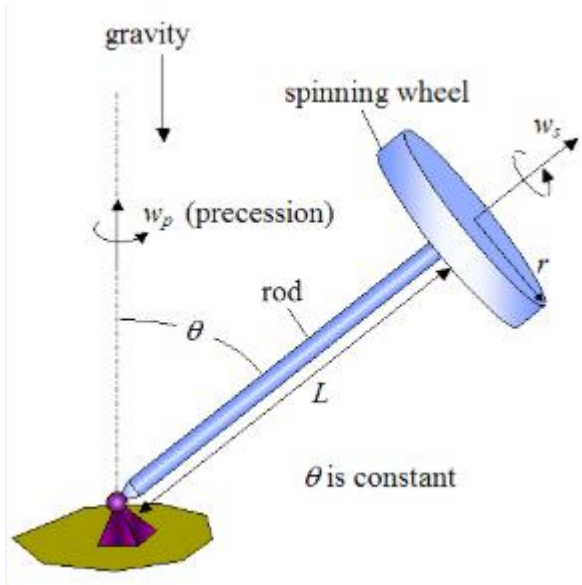
In general it is a combination of a physical quantity and a distance.

$$\vec{M} = \vec{F} \times \vec{r} \quad [\text{N} \cdot \text{m}] = [\text{kg} \cdot \text{m}^2 \cdot \text{s}^{-2}]$$



basic terms and quantities

precession of rigid rotating body: is the result of the gravitational force action and moment of the force (due to the rotation of the body)



similar effects act on rotating Earth

gyroscope (rotating flywheel)

https://www.youtube.com/watch?v=cquvA_lpEsA
(from timing 1:25)

basic terms and quantities

Momentum: is the product of the mass and velocity of an object. It is connected with the kinematic energy of the moving object.

It is in general a vector quantity.

$$\vec{p} = m\vec{v} \quad [\text{kg} \cdot \text{m} \cdot \text{s}^{-1}]$$

where m is mass [kg] and $\mathbf{v}$ is velocity [$\text{m} \cdot \text{s}^{-1}$].



$\neq$



basic terms and quantities

Pressure: Pressure is the amount of force acting per unit area.

$$p = \frac{|\vec{F}|}{a} = \frac{F}{a}$$

where F is the size of normal force [N] and a is the area of the surface on contact [m²],

unit is pascal [Pa] = [N/m²] = [kg·s⁻²·m⁻¹],

old unit was bar (1 bar = 100 000 Pa), (normal atmospheric/air pressure)

It is a scalar quantity.

Comment: In some physical applications, pressure can be connected with the direction of the acting force, we speak then about strain or stress

basic terms and quantities

Mechanical work:

In mechanics, a force is said to do work W if, when acting on a body, there is a displacement of the point of application in the direction of the force.

Its size is given by the product of force and distance.

Unit of work is joule $[J] = [N \cdot m] = [kg \cdot m^2 \cdot s^{-2}]$.

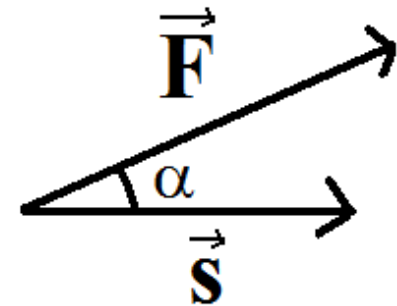
Mathematically it is a scalar product of force and distance (vectors):

$$W = \vec{F} \cdot \vec{s}$$

Size of scalar product is given:

$$W = |\vec{F}| |\vec{s}| \cos \alpha = F s \cos \alpha$$

It is a scalar quantity.



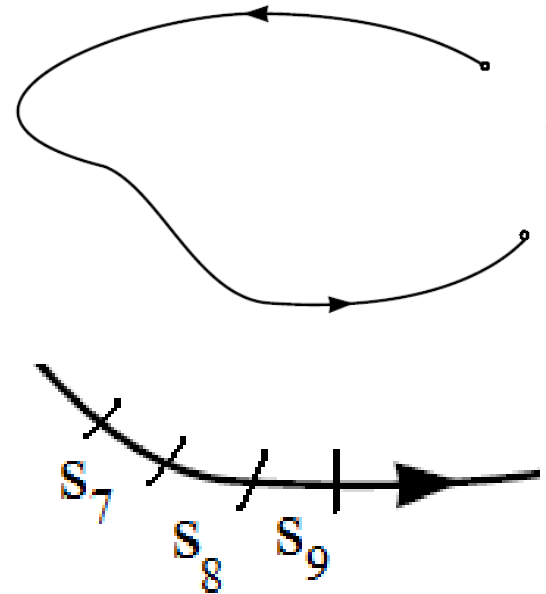
basic terms and quantities

Mechanical work:

But what to do, when the trajectory is not straight, but of irregular shape?

We can divide it into N small parts and evaluate work for each of them:

$$W = \sum_{i=1}^N \vec{F}_i \cdot \vec{s}_i$$



... and when the size of these small parts will be very small...?

$$W = \int_S \vec{F} \cdot d\vec{s} \quad \text{where } \vec{S} \text{ is the path and } d\vec{s} \text{ its differential (very small part).}$$

basic terms and quantities

Power:

Power is defined as the rate at which work is done upon an object. Like all rate quantities, power is a time-based quantity.

It is evaluated as the ration of work and time:

$$P = \frac{W}{t}$$

where W is work [J] and t is time [s].

Unit of power is watt [W] = [J·s⁻¹] = [kg·m²·s⁻³].

It is a scalar quantity.

Lecture 2: Mechanics

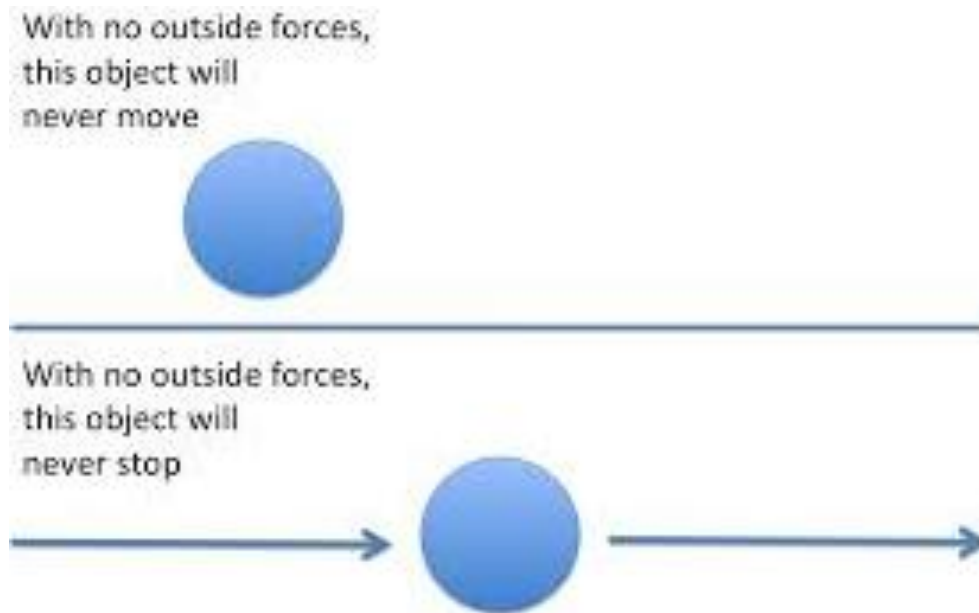
Content:

- basic terms and quantities
- velocity and acceleration
- force, moment of force, momentum
- work, power
- Newton's laws
- Kepler's laws
- Newton's gravitational law
- free fall, motion in gravitational field
- mechanical energy

Newton's **three laws of motion**:

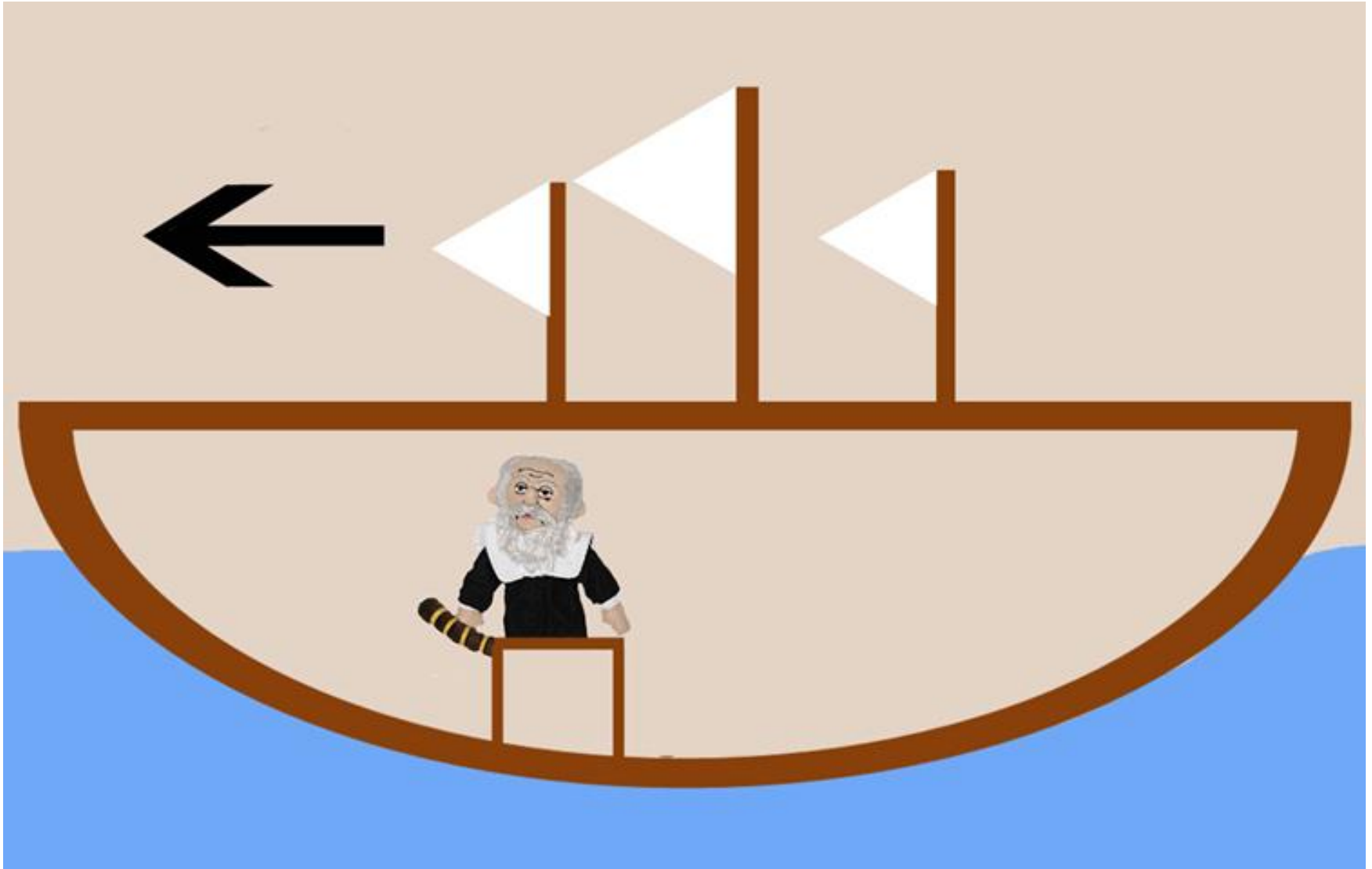
- I. Every object in a state of uniform motion tends to remain in that state of motion unless an external force is applied to it.

This we recognize as essentially Galileo's concept of inertia, and this is often termed simply the "**Law of Inertia**" (inertia is the tendency of matter to resist changes in its velocity).



Galileo's concept of inertia

(inertia is the tendency of matter to resist changes in its velocity)



Newton's **three laws of motion**:

II. The relationship between an object's mass m , its acceleration $\mathbf{a}$, and the applied force $\mathbf{F}$ is: $\mathbf{F} = m\mathbf{a}$ or $\vec{F} = m\vec{a}$
Acceleration and force are vectors; in this law, directions of the both vectors is the same. Simply the "**Law of Power**"

This is the **most powerful of Newton's three Laws**, because it allows quantitative calculations of dynamics: how do velocities change when forces are applied.

Notice the fundamental difference between Newton's 2nd Law and the dynamics of Aristotle: **according to Newton, a force causes only a change in velocity (an acceleration); it does not maintain the velocity as Aristotle held.**
Thus, **according to Aristotle there is only a velocity if there is a force**, but **according to Newton a force acts on it to cause an acceleration** (that is, a change in the velocity).

Newton's **three laws of motion**:

- II. The relationship between an object's mass m , its acceleration $\mathbf{a}$, and the applied force $\mathbf{F}$ is: $\mathbf{F} = m\mathbf{a}$ or $\vec{F} = m\vec{a}$
Acceleration and force are vectors; in this law, directions of the both vectors is the same.

Physical quantities and units:

$$[m] = \text{kg}$$

$$[\mathbf{a}] = \text{m} \cdot \text{s}^{-2} \quad (\text{change of velocity with respect to the time})$$

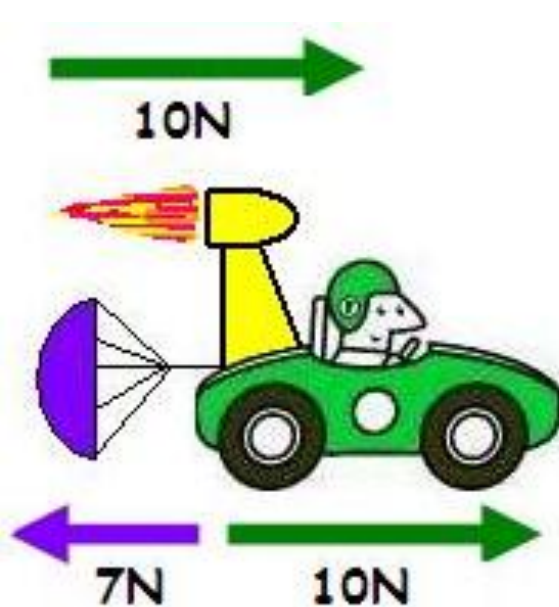
$$[\mathbf{F}] = \text{N} = \text{kg} \cdot \text{m} \cdot \text{s}^{-2}$$

Comment: velocity and acceleration defined by means of derivatives:

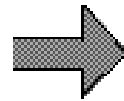
$$v = \frac{ds}{dt} = s' \quad [\text{m} \cdot \text{s}^{-1}] \quad a = \frac{dv}{dt} = \frac{d^2s}{dt^2} = v' = s'' = [\text{m} \cdot \text{s}^{-2}]$$

Newton's **three laws of motion**:

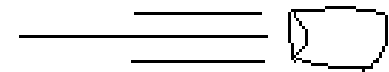
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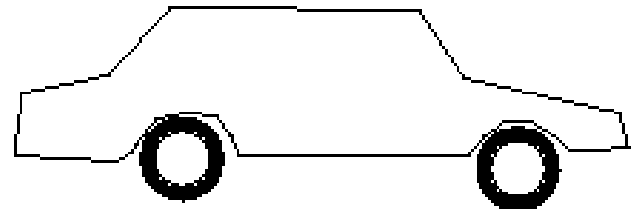
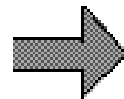
Same force



small mass: large acceleration



large mass: small acceleration



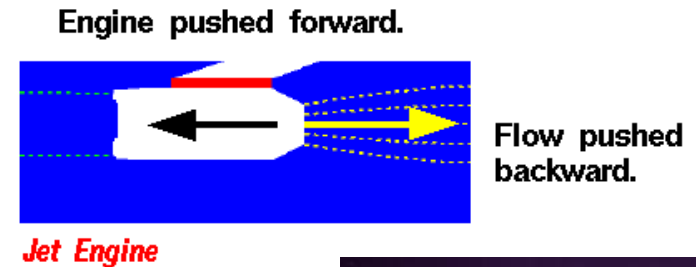
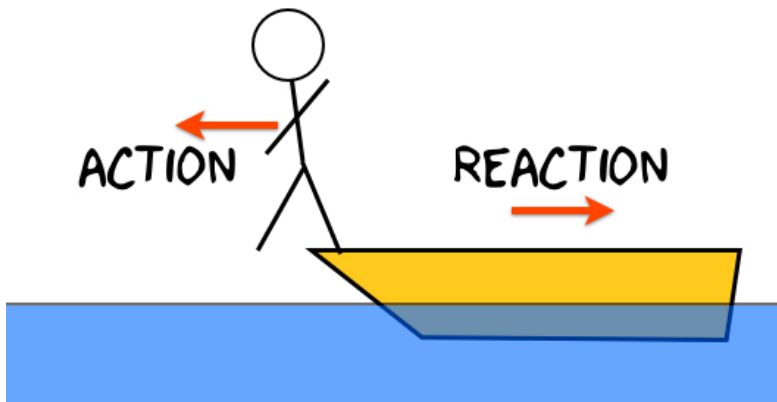
Force = mass x acceleration

Newton's **three laws of motion**:

III. For every action there is an equal and opposite reaction.

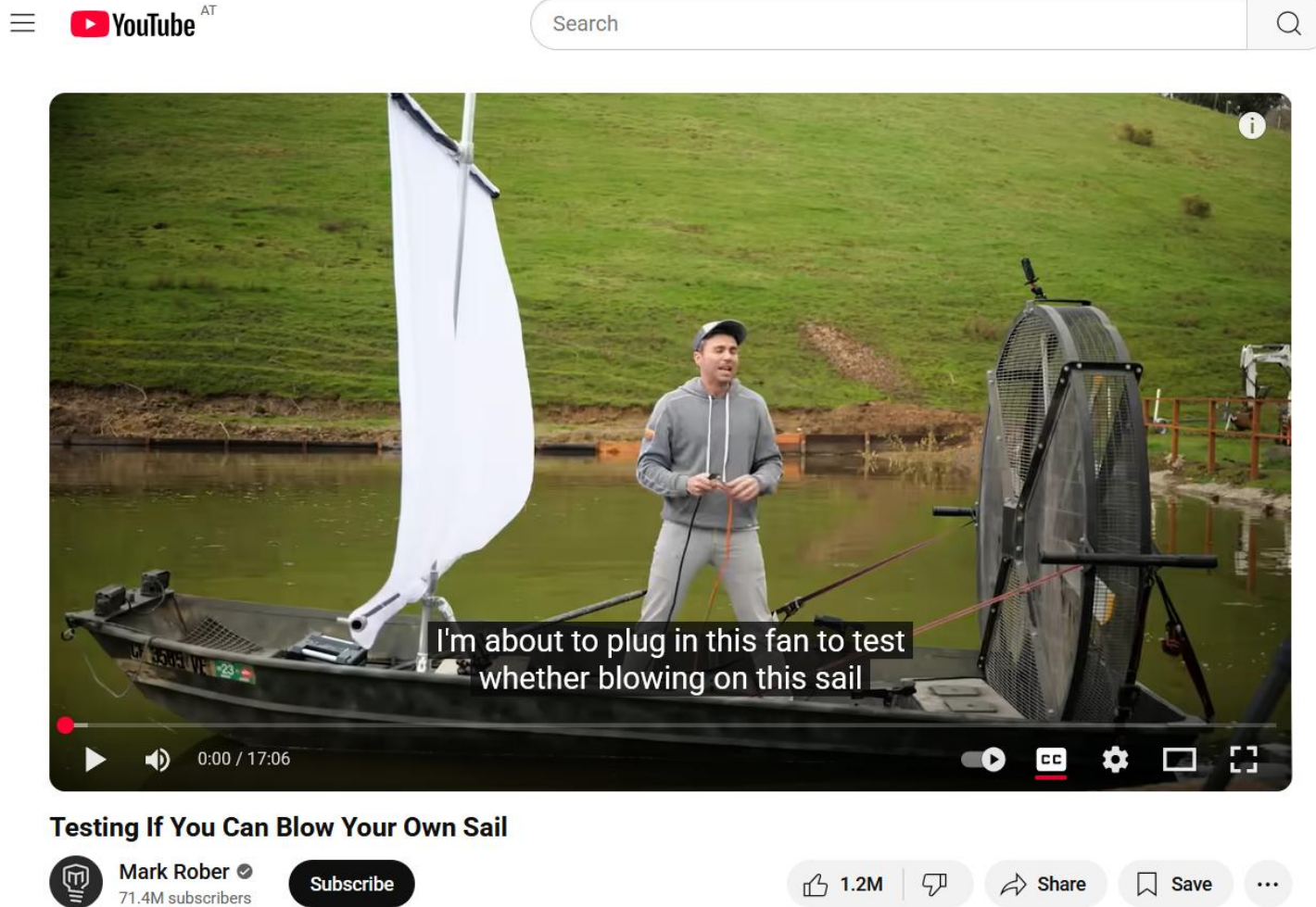
This is often termed simply the "**Law of action and reaction**".

This law is exemplified by what happens if we step off a boat onto the bank of a lake: as we move in the direction of the shore, the boat tends to move in the opposite direction (leaving us facedown in the water, if we aren't careful!).



Newton's **three laws of motion**:

III. "Law of action and reaction"



<https://www.youtube.com/watch?v=M7-h3FO-KKo>
(from timing 5:45)

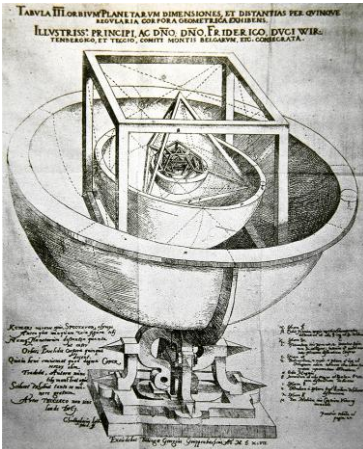
gravitation



g

Kepler's Work

- Tycho Brahe led a team which collected data on the position of the planets (1580-1600 with no telescopes).
- mathematician Johannes Kepler was hired by Brahe to analyze the data.

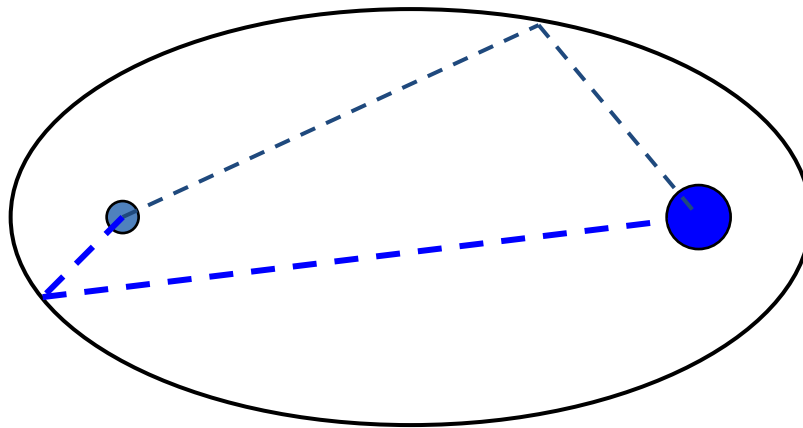


Johannes Kepler
1571 - 1630

- he took 20 years of data on position and relative distance.
- no calculus, no graph paper, no log tables.
- both Ptolemy and Copernicus were partly wrong.
- he **determined three laws of planetary motion** (1600-1630).

1. Kepler's First Law

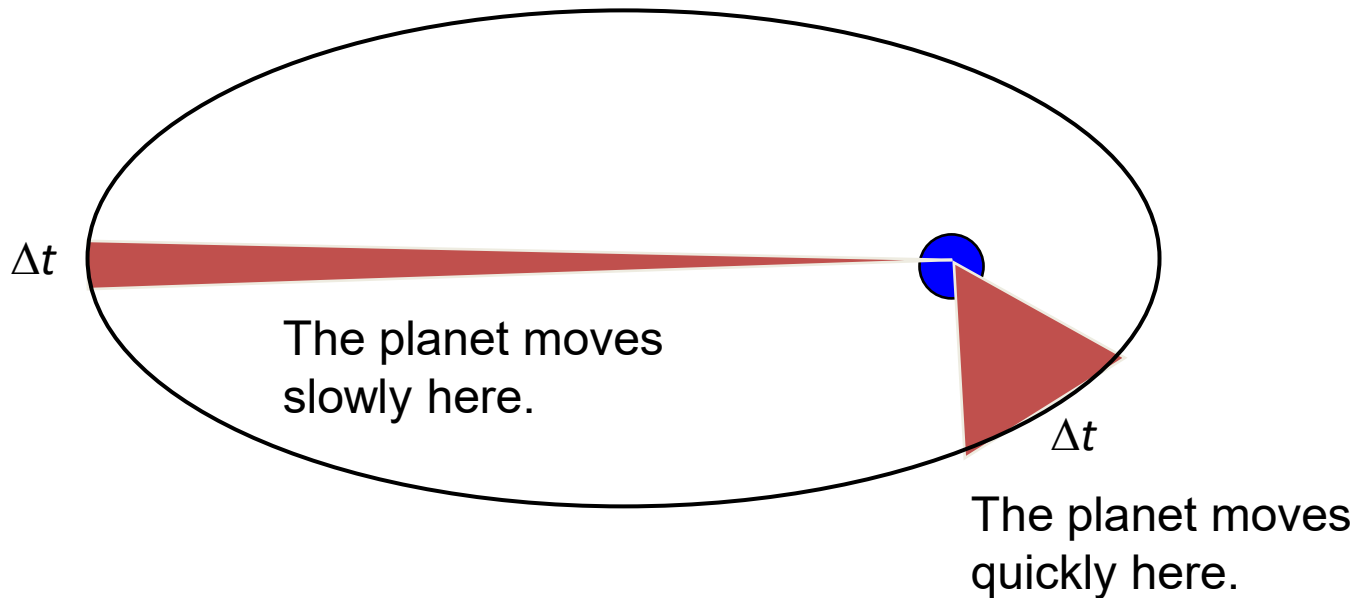
- The orbit of a planet is an ellipse with the sun at one focus.



A path connecting the two foci to the ellipse always has the same length.

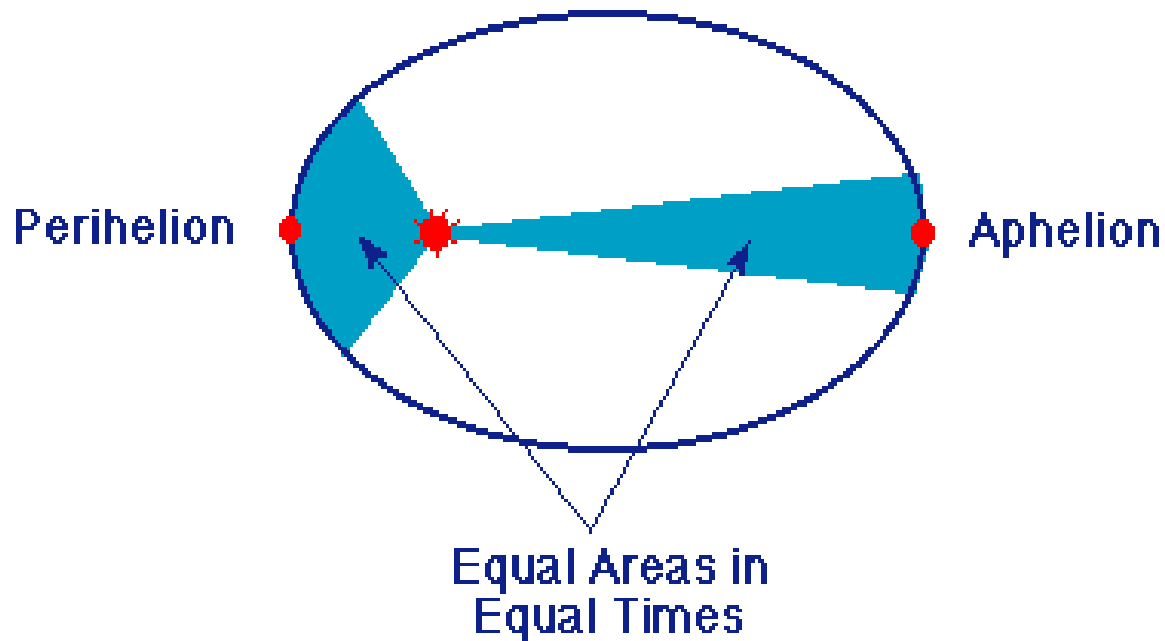
2. Kepler's Second Law

- The line joining a planet and the sun sweeps equal areas in equal time.



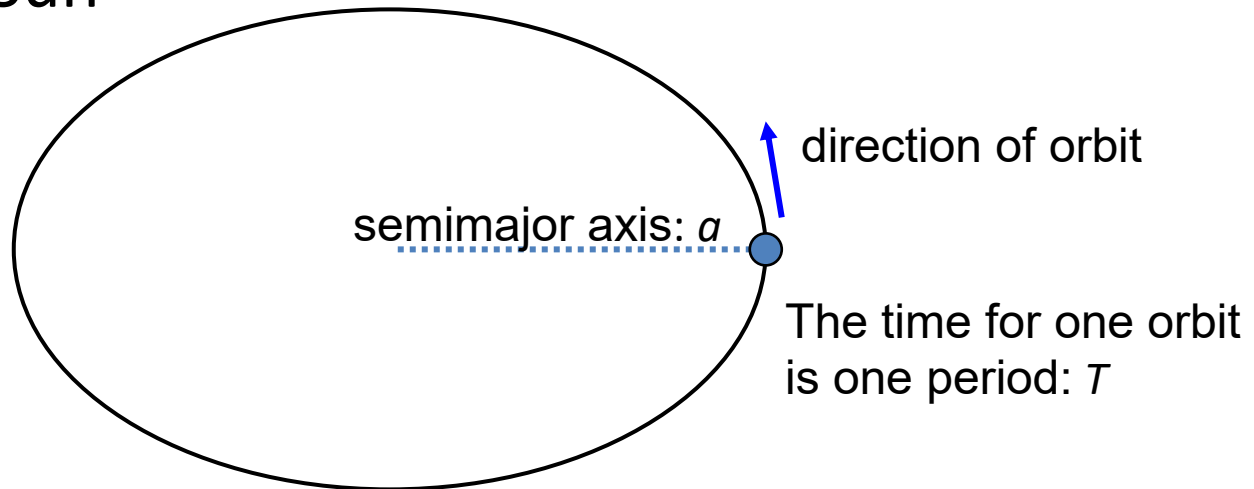
2. Kepler's Second Law

- The line joining a planet and the sun sweeps equal areas in equal time.



3. Kepler's Third Law

- The square of a planet's period is proportional to the cube of the length of the orbit's semimajor axis.
 - $T^2/a^3 = \text{constant}$
 - The constant is the same for all objects orbiting the Sun



Kepler's Third Law

Example: *planets Earth and Jupiter.*

Jupiter's period is **11.86 year** (11,86-times period of Earth),
semimajor axis (compared to Earth) is **5.2-times** larger.

So, it should be valid:

$$T^2/a^3 [\text{Earth}] = T^2/a^3 [\text{Jupiter}]$$

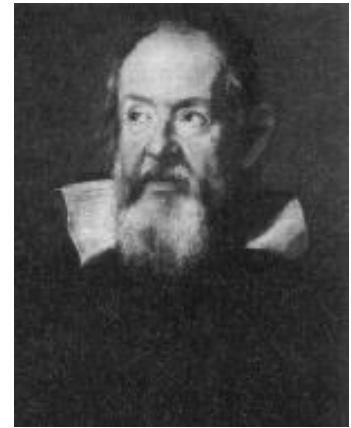
$$1^2/1^3 = (11.86)^2/(5.2)^3$$

$$(5.2)^3/1^3 = (11.86)^2/1^2$$

$$140.608 \approx 140.659$$



Work of Galileo Galilei



Galileo Galilei
(1564 - 1642)



- various contributions to the concept of modern science,
- mathematical derivations,
- astronomical observations,
- engineering experiments,
- free fall experiments
(velocity is independent from the body mass
- a contradiction to Aristotelian physics).

Very nice trial (Brian Cox, vacuum chamber):

<https://www.youtube.com/watch?v=E43-CfukEgs&feature=youtu.be>

Experiment on the Moon (Apollo 15):

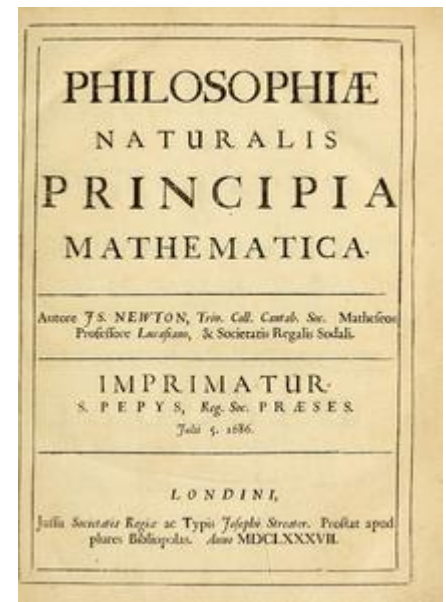
https://commons.wikimedia.org/wiki/File:Apollo_15_feather_and_hammer_drop.ogv

Newton's Work

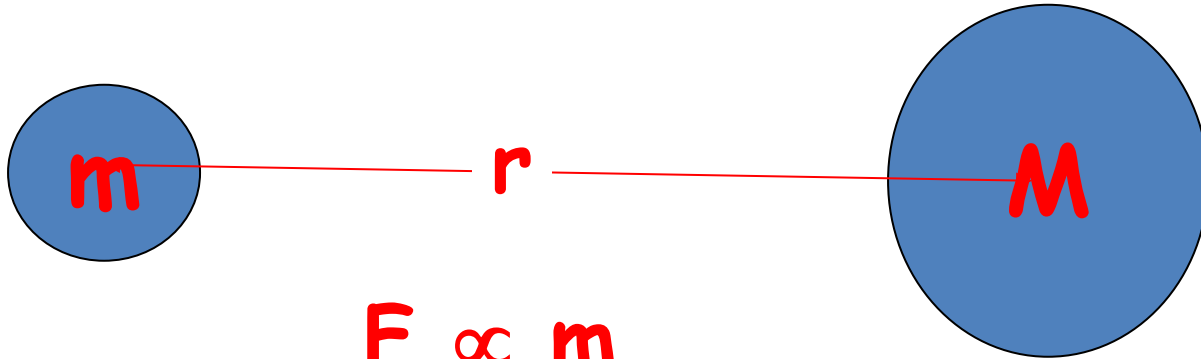
- laws of motion
- universal law of gravity
- mathematical derivation of Kepler's laws
- introduction of calculus (derivatives)
- most important work:
Philosophiæ Naturalis Principia Mathematica
("Mathematical Principles of Natural
Philosophy"), first published 5 July 1687
(later edited versions: 1713 and 1726)
- he spent the second half of his life in Royal mint



Isaac Newton
1643 - 1727



Universal law of gravity



$$F \propto m$$

$$F \propto M$$

$$F \propto \frac{1}{r^2}$$

combine:

$$F \propto \frac{mM}{r^2}$$

Proportionality constant:

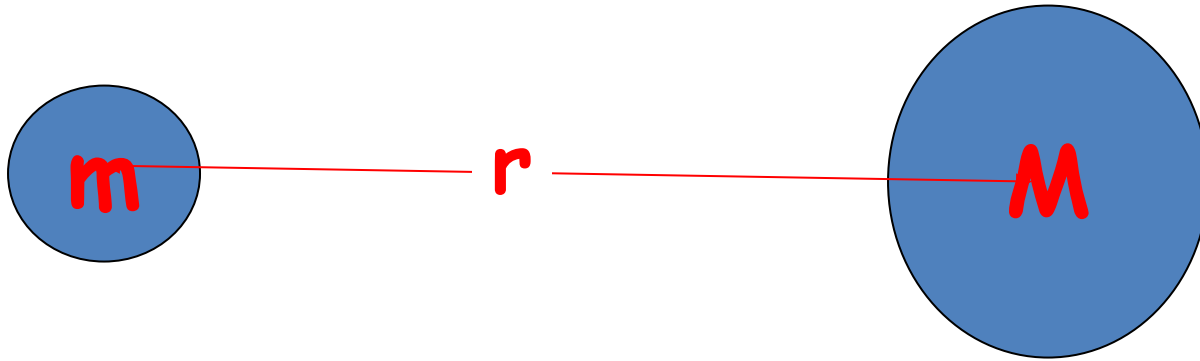
"Newton's Constant"

$$F = G \frac{mM}{r^2}$$

Comment:

This property ($1/r^2$) was emphasized before Newton by Robert Hook and Ismaël Boulliau.

Universal law of gravity



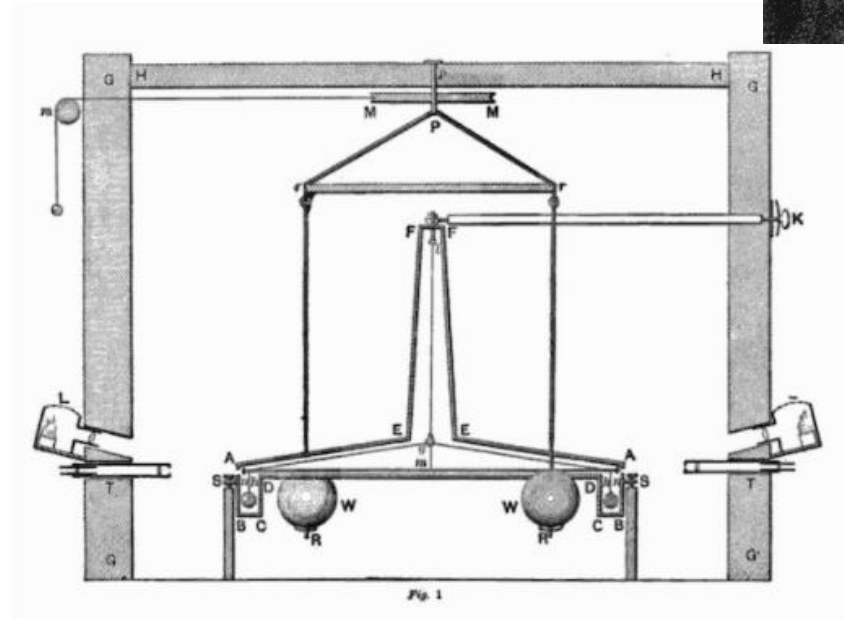
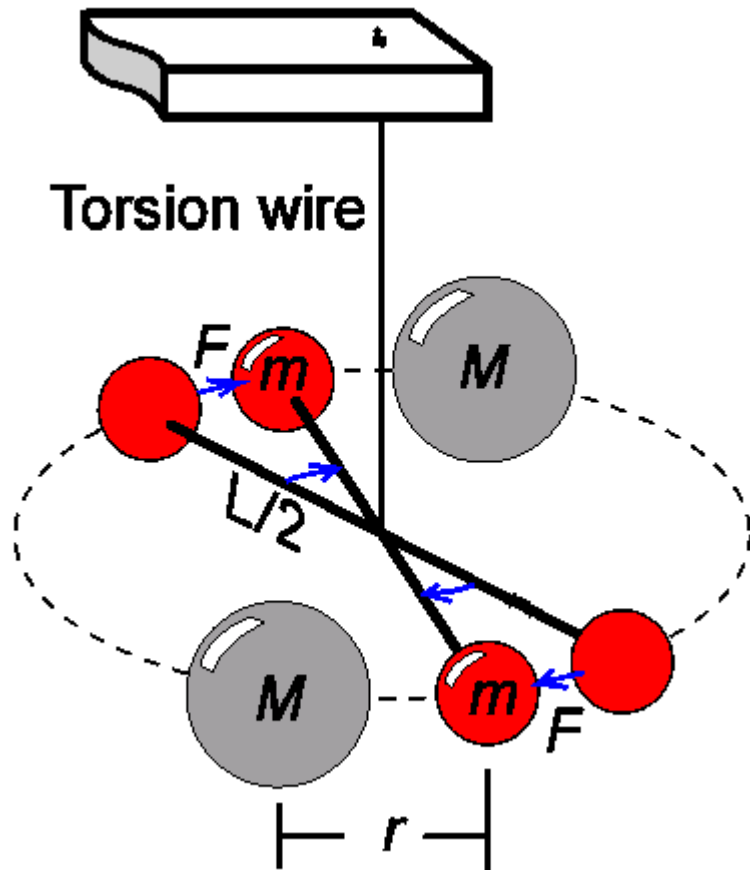
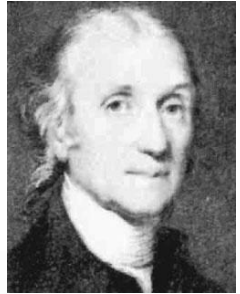
Newton's law of universal gravitation states that any two bodies in the universe attract each other with a force that is directly proportional to the product of their masses (m, M) and inversely proportional to the square of the distance between them (r).

$$|\vec{F}_G| = G \frac{mM}{r^2}$$

$$G = \kappa = 6.674 \cdot 10^{-11} \left[\text{m}^3 \cdot \text{kg}^{-1} \cdot \text{s}^{-2} \right]$$

G – is universal gravitational constant
(estimated for the first time by H. Cavendish in 1797-1798)

Measuring gravity force between “ordinary-sized” objects is not easy..



Cavendish's value:

$$G = 6.74 \cdot 10^{-11} \left[\text{m}^3 \cdot \text{kg}^{-1} \cdot \text{s}^{-2} \right]$$

$$\text{today: } G = 6.674\text{XX} \cdot 10^{-11} \left[\text{m}^3 \cdot \text{kg}^{-1} \cdot \text{s}^{-2} \right]$$

gravitational acceleration (g):

Newton's gravity law:

$$F = G \frac{mM}{r^2}$$

2. Newton's
motion law:

$$F = mg \Rightarrow g = \frac{F}{m} \Rightarrow g = G \frac{M}{r^2} \quad [\text{m} \cdot \text{s}^{-2}]$$

Value of g?

In our country approx. $9.81 \text{ m} \cdot \text{s}^{-2}$ (rounded $10 \text{ m} \cdot \text{s}^{-2}$).

It is not a constant! Its value is influenced by many factors (rotation of Earth, distance from Earth center, large masses on the surface or below it).

But in the same place on the Earth it acts on falling object independently on their mass (in vacuum).

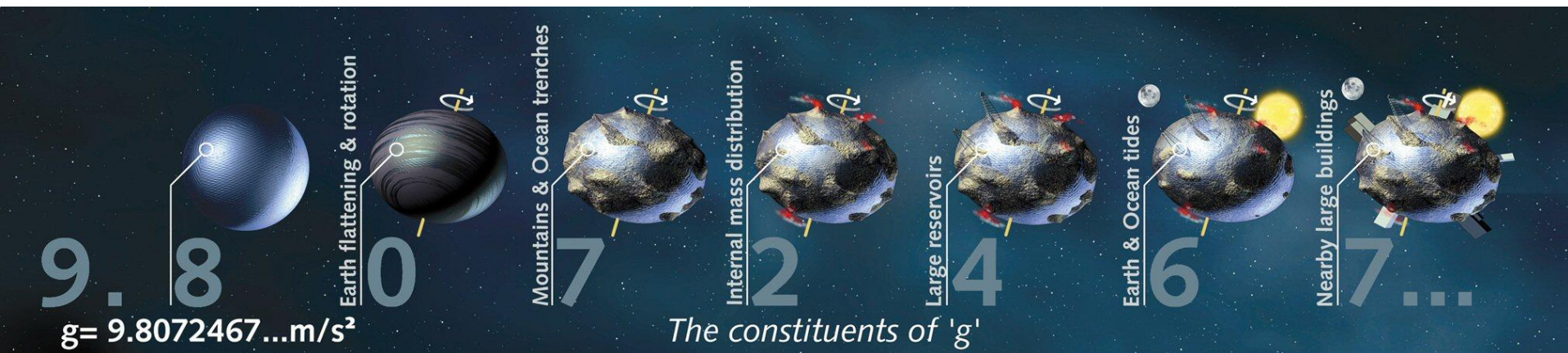
gravitational acceleration (g):

Estimation of the g value for our Earth:

(mass of the Earth $\sim 5.97 \cdot 10^{24}$ kg; radius ~ 6371000 m,
 $G \sim 6.674 \cdot 10^{-11}$ N·m²/kg²)

$$g = G \frac{M}{r^2} \approx 9.8 \text{ [m} \cdot \text{s}^{-2}]$$

In one point at the Earth surface g value is constant (independent from the mass), but it changes with the change of position (!)

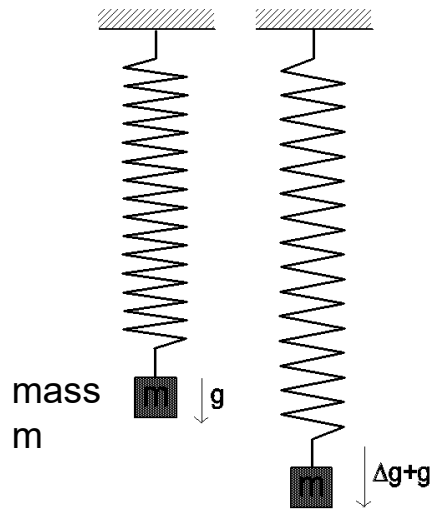


gravitational acceleration (g):

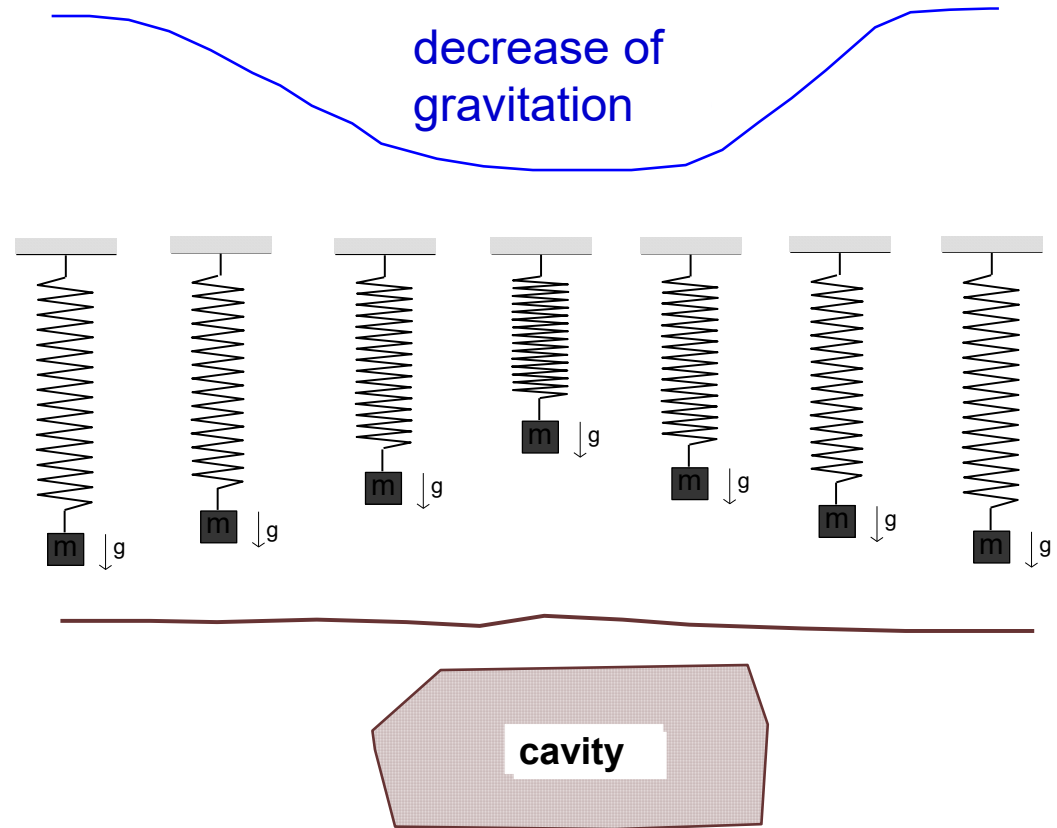
relative gravimeter:



measuring principle:



scheme of a spring
relative gravimeter



decrease of measured gravitational acceleration
above a cavity

free fall and vertical throw

free fall

exact derivation
is in the appendix
of this lecture

$$s = v_0 t + \frac{1}{2} g t^2$$

When we take $v_0 = 0$ (velocity for time $t = 0$), we get the well-known formula:

$$s = \frac{1}{2} g t^2$$

$$v = v_0 + g t$$

change of the velocity

Example:

$$t_1 = 1 \text{ sec} \Rightarrow s_1 = 0.5 g t_1^2 = 5 \cdot 1 = 5 \text{ m}$$

$$t_2 = 2 \text{ sec} \Rightarrow s_2 = 0.5 g t_2^2 = 5 \cdot 4 = 20 \text{ m}$$

$$t_3 = 3 \text{ sec} \Rightarrow s_3 = 0.5 g t_3^2 = 5 \cdot 9 = 45 \text{ m}$$

$$t_4 = 4 \text{ sec} \Rightarrow s_4 = 0.5 g t_4^2 = 5 \cdot 16 = 80 \text{ m}$$



free fall

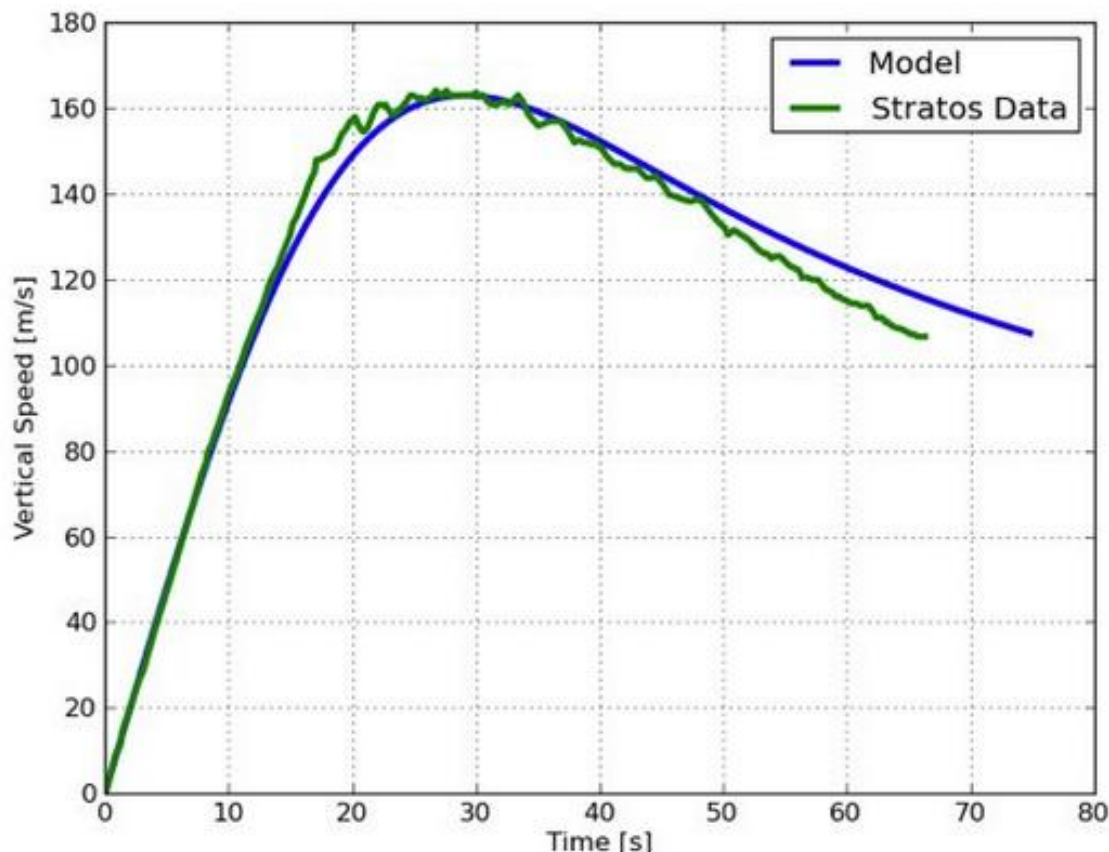
$$s = \frac{1}{2}gt^2 \Rightarrow t = \sqrt{\frac{2s}{g}}$$

Example: jump of Felix Baumgartner (2012)

height: 38 969 m

time: 4 min 20 sec.

Can we check it by means of free fall formula?



Air resistance:

$$|\vec{F}_{\text{AIR}}| = -k|\vec{v}|^2$$

where k is the air resistance coefficient
 $k = 0.24 \text{ [kg/m]}$

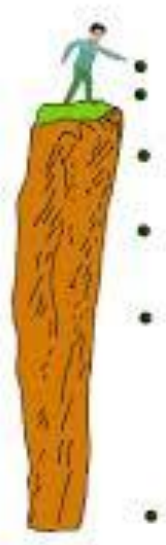
vertical throw

$$s = v_0 t - \frac{1}{2} g t^2$$

$$v = v_0 - g t$$

change of the velocity

The motion is damped, until $v = 0$ and from this moment a free fall starts.



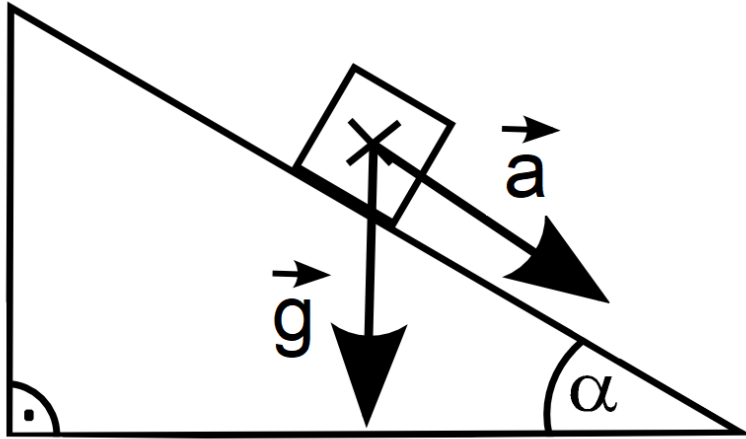
free fall



vertical throw

motion along inclined plane

(without friction)



$$\mathbf{F}_G = m\mathbf{g}$$

from the figure it follows:

$$\begin{aligned}\mathbf{a} &= \mathbf{F}_G/m = (\mathbf{F}_G \sin\alpha)/m = \\ &= (m \mathbf{g} \sin\alpha)/m = \mathbf{g} \sin\alpha\end{aligned}$$

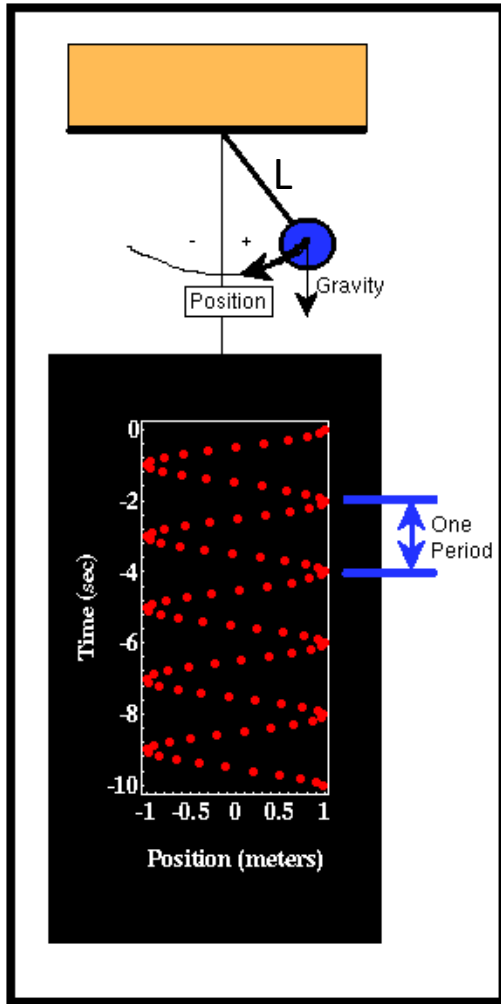
$$\mathbf{a} = \mathbf{g} \sin \alpha$$

From this derivation and also from experiments it follows that the accelerated motion along inclined plane is also **independent from the body mass**.

Valid also in models of water transport on slopes in hydrology.

Comment: This is not valid in a case of rolling object.

mathematical pendulum – period



Period of a mathematic pendulum (T):

Is also independent on the mass of the object, it is a function of the length L and gravitational acceleration g :

$$T = 2\pi \sqrt{\frac{L}{g}} \quad [\text{s}]$$

Walter Lewin – lecture at MIT
(video),

$$L = 5.21 \text{ m}, g = 9.8 \text{ m/s}^2,$$

$$T = 4.58 \text{ s}$$

http://www.youtube.com/watch?v=KXys_mymMKA

mathematical pendulum – derivation of the period

The main force, causing the movement

is connected with the action of gravity force: $F = F_G \sin \alpha$ (1)

From the main triangle (picture) it is valid ($\alpha < 5^\circ$): $\sin \alpha = y/L$

And we use also following relations: $F_G = mg$, $F = ma$

Starting with Eq. (1): $F = F_G \sin \alpha$

$$ma = mg \sin \alpha$$

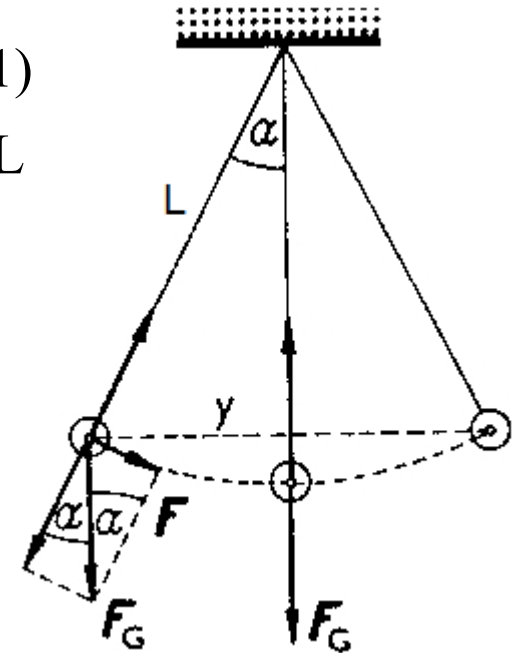
$$a = g \frac{y}{L} \quad (2)$$

We know from the rotational motion: $a = \omega^2 y$, $\omega = 2\pi/T$

where: ω is the angular velocity and T period.

Setting this into Eq. (2) we obtain:

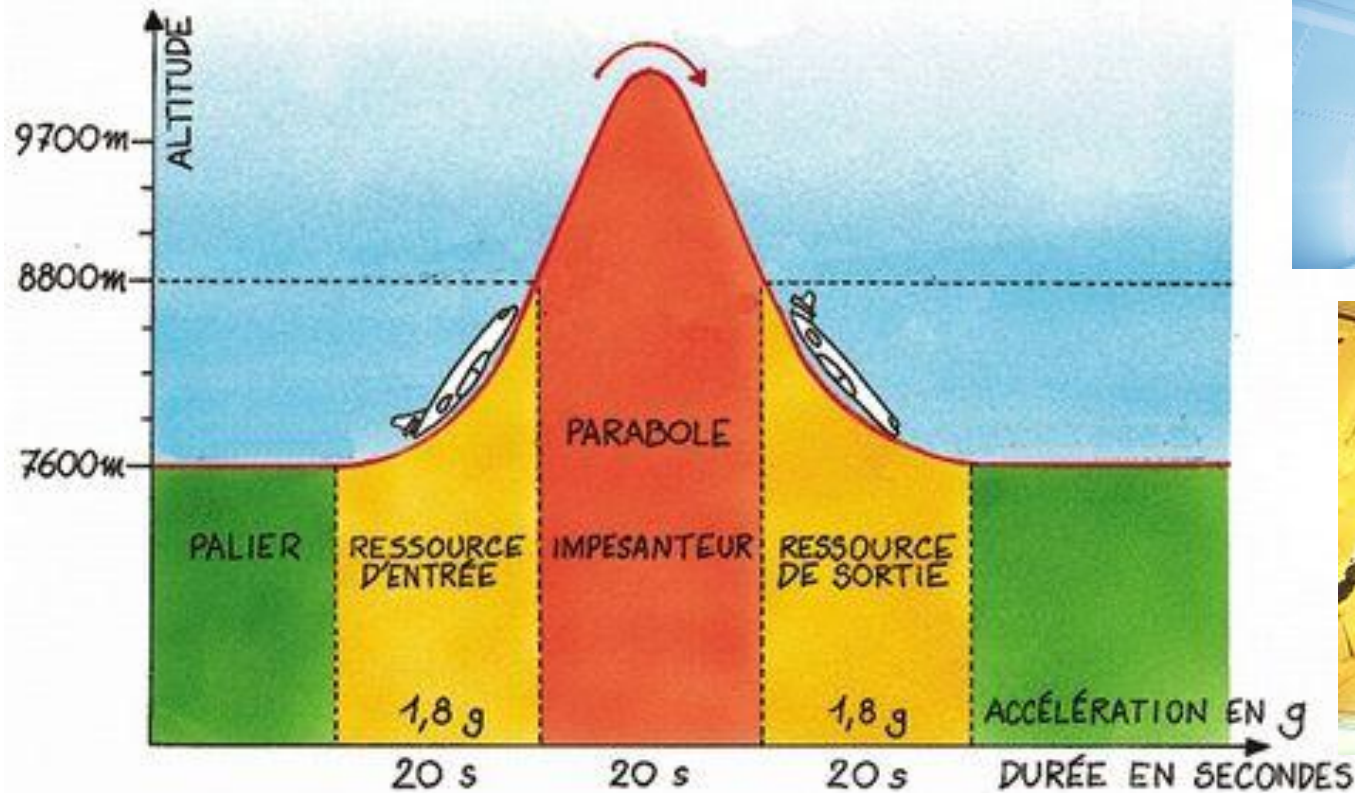
$$\omega^2 y = g \frac{y}{L}$$
$$\left(\frac{2\pi}{T} \right)^2 = \frac{g}{L} \Rightarrow T = 2\pi \sqrt{\frac{L}{g}}$$





simulation of weightless stage (in an aircraft) – zero G parable

(used for astronauts training)



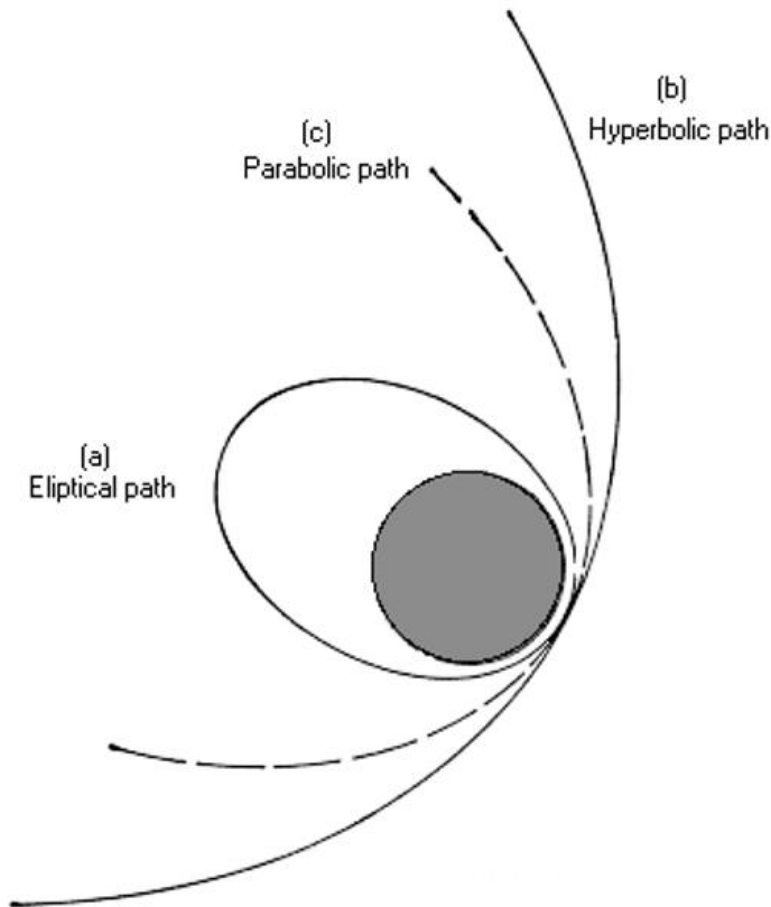
used also for commercial purposes:

<http://www.gozerog.com>

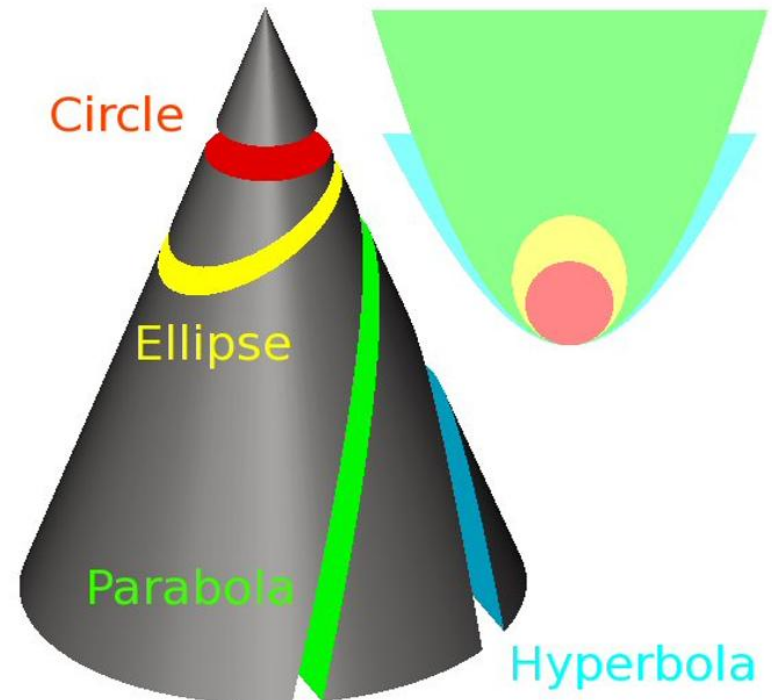
trajectories of movement in gravitational field

In larger scales (planetary scales) and in a central gravitational field (main source is only one body) – shapes of trajectories are conic sections (circle, ellipse, parabola or hyperbola).

This shape depends on a balance between gravity, mass, velocity and distance.



conic sections:

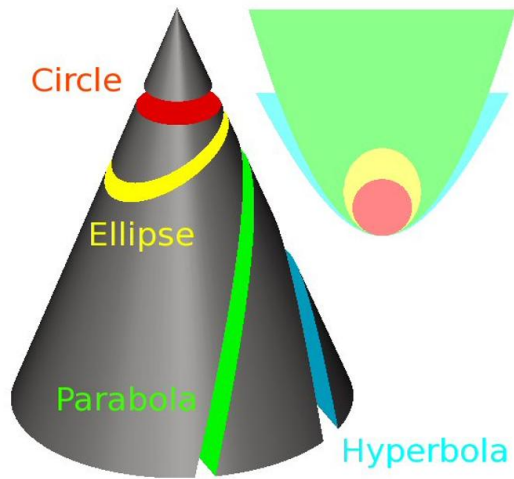


trajectories of movement in gravitational field

In larger scales (planetary scales) and in a central gravitational field (main source is only one body) – shapes of trajectories are conic sections (circle, ellipse, parabola or hyperbola).

This shape depends on a balance between gravity, mass, velocity and distance.

conic sections:



$$f = (a - c)/a$$

a – major semi-axis, c – minor semi-axis

Planetary Orbital Eccentricities

Planet	flattening . f	Shape Description
Mercury	0.206	Most elliptical planet
Venus	0.007	Most circular planet
Earth	0.017	Nearly circular
Mars	0.093	Moderately elliptical
Jupiter	0.048	Nearly circular
Saturn	0.054	Nearly circular
Uranus	0.047	Nearly circular
Neptune	0.009	Nearly circular

trajectories of movement in gravitational field

Also in smaller scales (surface of the Earth) shapes of trajectories can reach some of the conic sections:

time-laps records (a basketball being thrown):

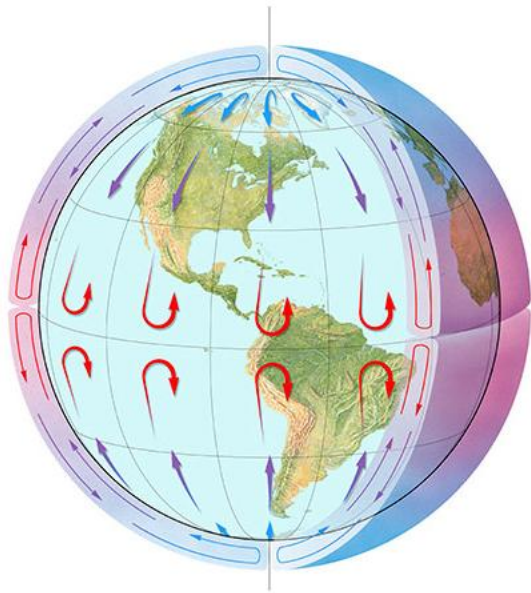


Coriolis force (acceleration)

acting due to the rotation of the Earth

(acting on objects moving in N-S and S-N direction)

This force (acceleration) plays an important role in global meteorology, in hydrology and also in aeronautics.



cyclone
(large rotating air mass)



meandering of rivers

tricks on equator (Quito, Ecuador):

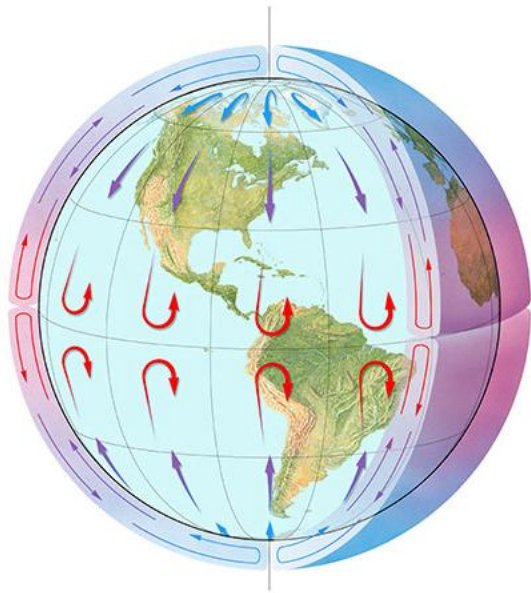
<https://www.youtube.com/watch?v=4IIVfoDuVlw>

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Coriolis effect on equator (Quito, Ecuador):

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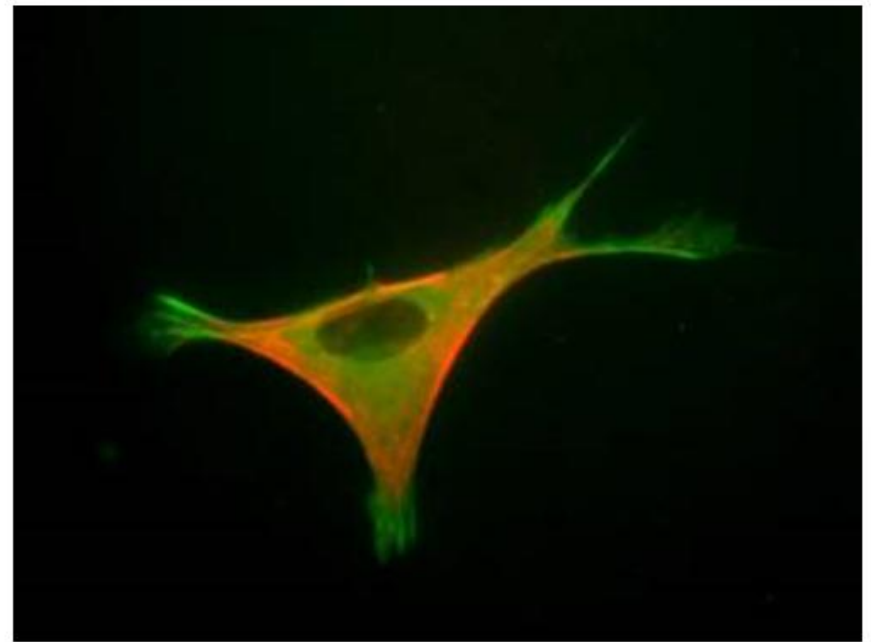
<https://www.youtube.com/watch?v=M7-h3FO-KKo>

(Mark Rober video, from timing 12:10)

Example: gravity and biology

Weightless stage – and its influence on muscles function

Muscle cells have unique ways of detecting mechanical stress. Scientists think that the lack of mechanical stress on cells from gravity may decrease tension in the cell membrane and affect the expression of key proteins and genes, ultimately leading to muscle atrophy.



The type of mouse cell used in the Cell Mechanosensing investigation.

Credits: JAXA/Nagoya University

Lecture 2: Mechanics

Content:

- mechanics - basic terms and quantities
- velocity and acceleration
- force, moment of force, momentum
- work, power
- Newton's laws
- Kepler's laws
- Newton's gravitational law
- free fall, motion in gravitational field
- **mechanical energy**

back to basic terms and quantities

Mechanical energy:

It is the energy associated with the motion and position of an object:

- kinetic energy (E_k),
- potential energy (E_p).

In so called conservative fields the sum of potential energy and kinetic energy is constant.

Additional energies in mechanics:

- energy of rotation body,
- elastic energy.

Unit of energy is identical with the unit of mechanical work (joule) [J].

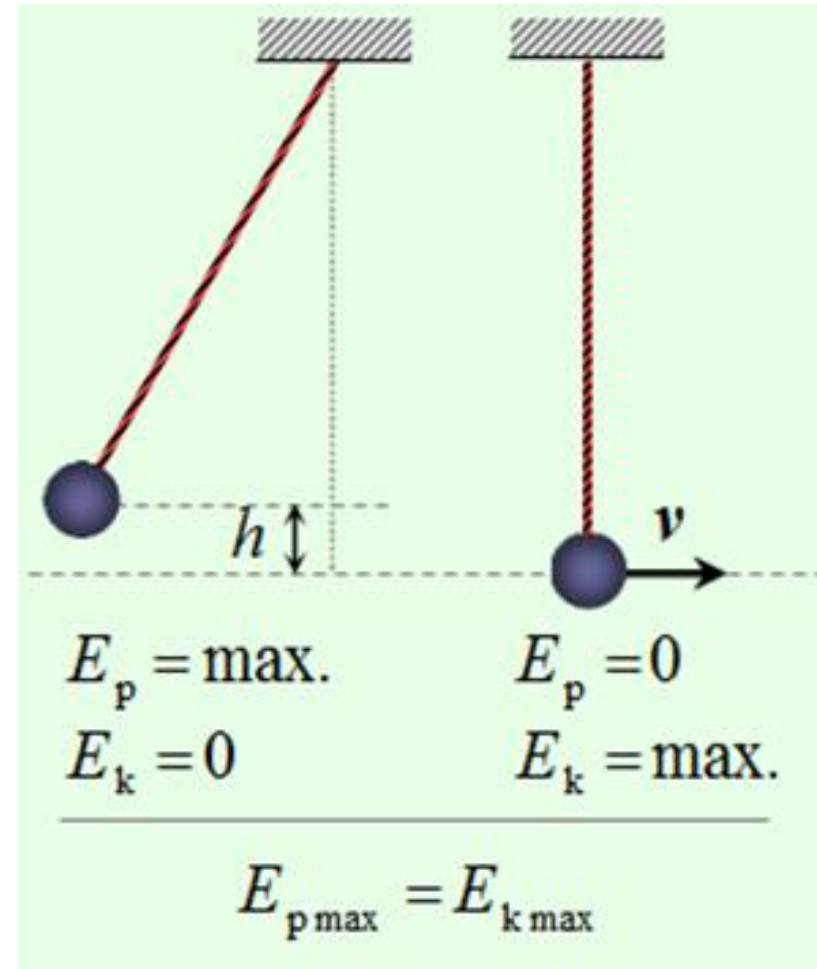
back to basic terms and quantities

Mechanical energy:

- kinetic energy (E_k),
- potential energy (E_p).

In so called conservative fields the sum of potential energy and kinetic energy is constant (only the form of the energy is changed).

This property is connected with the so called:
law of conservation of energy.



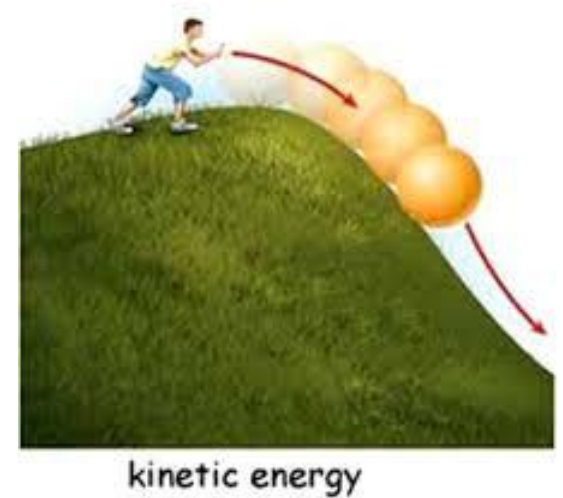
basic terms and quantities

Mechanical energy:

kinetic energy (E_k) - the energy that it possesses due to its motion.

$$E_k = \frac{1}{2}mv^2$$

where m is the mass [kg] and
 v the velocity [$\text{m}\cdot\text{s}^{-1}$].



This is valid in classical mechanics.

In relativistic mechanics, this is a good approximation only when v is much less than the speed of light.

kinetic energy - derivation:

Energy is connected with work: $\Delta E = F \Delta s$

for: $F = ma$ $\Delta s \approx v \Delta t$

$$\Delta E \approx m a v \Delta t$$

and: $a \Delta t = \Delta v$

is valid: $\Delta E \approx m v \Delta v$

velocity: $v = \frac{\Delta s}{\Delta t}$

acceleration: $a = \frac{\Delta v}{\Delta t}$

kinetic energy - derivation:

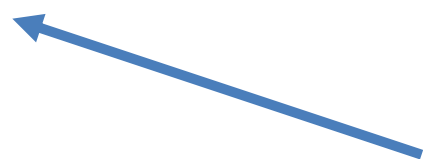
Energy is connected with work: $\Delta E = F \Delta s$

for: $F = ma$ $\Delta s \approx v \Delta t$

is valid: $\Delta E \approx m a v \Delta t$

and: $a \Delta t = \Delta v$

is valid: $\Delta E \approx m v \Delta v$



But where we got the $\frac{1}{2}$ in the result?:

$$\Delta(v^2) = (v + \Delta v)^2 - v^2 = 2v\Delta v + (\Delta v)^2 \approx 2v\Delta v \Rightarrow v\Delta v = \frac{1}{2} \Delta(v^2)$$

We have ignored the term $(\Delta v)^2$, because it is a very small number (e.g. $0.01^2 = 0.0001$).

Final expression:

$$\Delta E \approx \frac{1}{2} m \Delta(v^2) = \Delta\left(\frac{1}{2} m v^2\right)$$

Generally:

$$E_k = \frac{1}{2} m v^2$$

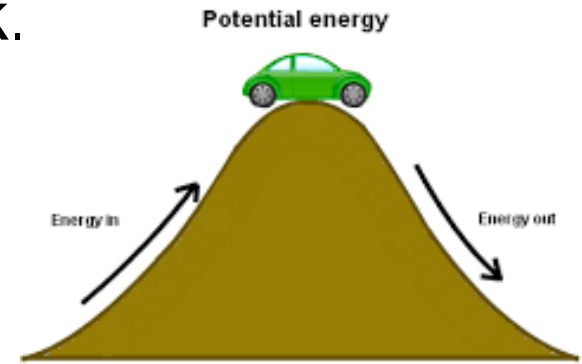
basic terms and quantities

Mechanical energy:

potential energy (E_p) - the energy that an object has due to its position in a force field (mostly gravitational field) – it is connected with the mechanical work.

$$E_p = mgh$$

where m is the mass [kg],
 g the gravitational acceleration [$\text{m}\cdot\text{s}^{-2}$]
and h the height [m].



The change of potential energy is dependent only from the height difference between two points and not from the trajectory of the motion between them.

basic terms and quantities

Moment of inertia: is a measure of an object's resistance to changes in the rotation direction.

For a point mass it can be expressed as:

$$I = mr^2 \quad [\text{kg} \cdot \text{m}^2]$$

where r is the distance of the point mass from the rotation axis.

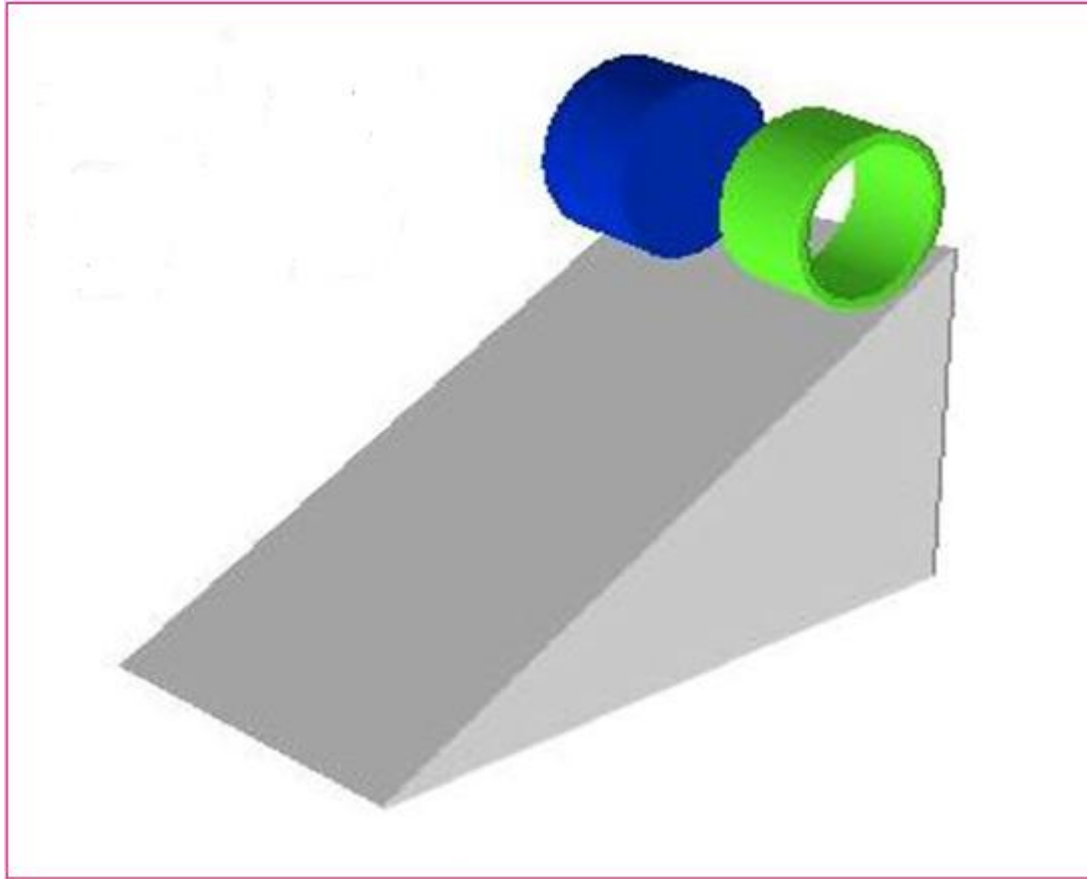
Energy of rotating body:

$$E_r = \frac{1}{2} I \omega^2$$

where ω is the size of angular velocity.

Energy of rotating body:

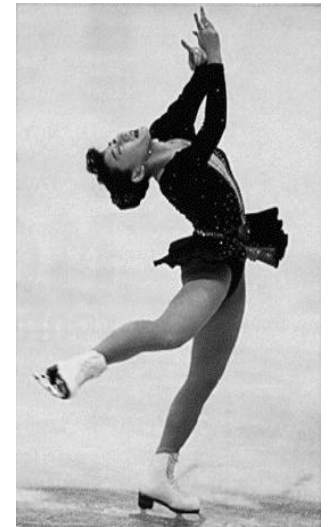
Problem solution: two bodies (blue and green)



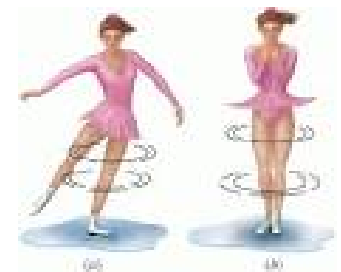
$$E_r = \frac{1}{2} I \omega^2$$

$$I = mr^2$$

Hint: pirouette
of a figure-skater



- a) identical masses
- b) different distances of masses from the centre of body rotation
- c) which one will move faster (will have higher ω)?



Energy of rotating body:



$$E_r = \frac{1}{2} I \omega^2$$

$$I = mr^2$$

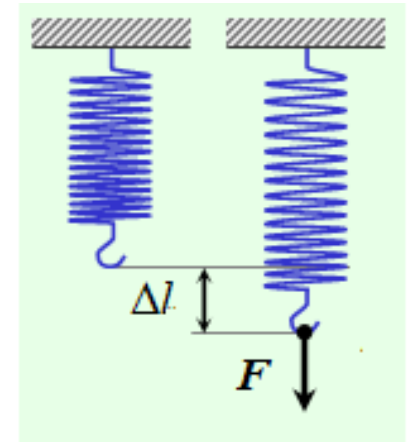
rocky landslide in Vyhne (central Slovakia)
- so called gravitational separation
(largest pieces at the bottom of the slope)

basic terms and quantities

Mechanical energy:

Elastic energy (E_{El}) - is the potential mechanical energy stored in the configuration of a material or physical system as work is performed to distort its volume or shape. Elastic energy occurs when objects are compressed and stretched, or generally deformed in any manner.

$$E_{El} = \frac{1}{2} k \Delta l^2$$



where k is so called spring (Hook) constant and Δl is the length change of the spring.

Appendix 1 – derivation of the free-fall equations (using 1D integration)

free fall – basic equations (1/2)

From the definition of velocity and acceleration it follows:

$$g = \frac{\partial^2 s}{\partial t^2}$$

Integrating this equation with respect to t , we get:

$$\int g dt = \int \left[\frac{\partial^2 s}{\partial t^2} \right] dt$$

$$g \int dt = \frac{\partial s}{\partial t} + c_1$$

$$gt + c_2 = \frac{\partial s}{\partial t} + c_1$$

$$gt = v + c_3$$

$$v = gt + v_0$$

Accepting the original condition that for the time $t = 0$ the initial velocity of the object at some level z_0 is v_0 , we get: $c_3 = -v_0$.

free fall – basic equations (2/2)

$$v = gt + v_0$$

In further step we integrate this equation again with respect to t:

$$\int v dt = \int [gt + v_0] dt$$

$$s + c_4 = \int gtdt + \int v_0 dt$$

$$s + c_4 = g \int t dt + v_0 \int dt$$

$$s + c_4 = g \frac{t^2}{2} + c_5 + v_0 t + c_6$$

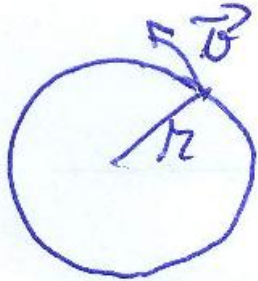
$$s = g \frac{t^2}{2} + v_0 t + c_7$$

$$s = \frac{1}{2}gt^2 + v_0t + z_0$$

Accepting the original condition that for the time $t = 0$ the position of the object is the level z_0 we get: $c_7 = z_0$.

Appendix 2 –
derivation of the formula for tangential velocity
(rotational motion)

rotational motion – tangential velocity (1/1)



angular velocity
 $|\vec{\omega}| = \frac{\Delta\phi}{\Delta t}$

tangential velocity
for constant velocity (speed)

$$\omega = \frac{\phi}{T} = \frac{2\pi}{T} \Rightarrow T = \frac{2\pi}{\omega} \quad \text{Eq. 1}$$

tangential component of velocity:

$$|\vec{v}_t| = v_t = v = \frac{L}{T} = \frac{2\pi r}{T} \Rightarrow T = \frac{2\pi r}{v} \quad \text{Eq. 2}$$

comparing
eqs. 1 and 2

$$\frac{2\pi}{\omega} = \frac{2\pi r}{v} \quad | : 2\pi$$
$$| \cdot \omega | \cdot v$$

$$v = \omega r$$

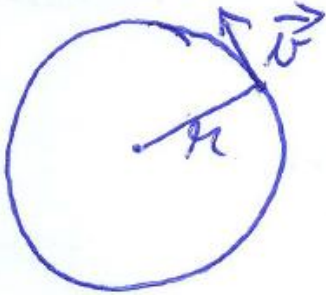
$$|\vec{v}_t| = |\vec{\omega}| r$$

T - period
L - circumference

Appendix 3 –
derivation of the formula for the normal component
of the centripetal/centrifugal acceleration
(rotational motion)

rotational motion – centripetal/centrifugal acceleration (1/3)

During a motion along a circle (circular motion), velocity $\vec{v}$ has only tangential component $|\vec{v}| = |\vec{v}_t| = v$, its normal component has zero size $|\vec{v}_n| = 0$.



Velocity $\vec{v}$ causes the change of the position. Period T can be easily calculated by the circumference of the circle ($2\pi r$):

$$T = \frac{2\pi r}{v} \quad (1)$$

Eq. 1

On the other hand, centripetal acceleration $\vec{a}$ has only the normal component (which forces the motion to change the direction – direction of $\vec{v}$), $|\vec{a}| = |\vec{a}_n| = a$, the tangential component has zero size $|\vec{a}_t| = 0$.

T – period, v – velocity, a – acceleration,
 r – radius of the circle

rotational motion – centripetal/centrifugal acceleration (2/3)

Acting of the centripetal acceleration can be in a symbolic way plotted by means of an analogous circle – where an analogous formula is valid:



$$T = \frac{2\pi |\omega|}{|a|} \quad (2)$$

Eq. 2

From the equality of these two periods it is valid:

comparing
eqs. 1 and 2

$$\begin{aligned} \frac{2\pi r}{v} &= \frac{2\pi v}{a} & | \cdot v & \quad | \cdot a \\ 2\pi r a &= 2\pi v v & | : 2\pi & \\ a r &= v^2 & | : r & \\ a &= \frac{v^2}{r} \end{aligned}$$

rotational motion – centripetal/centrifugal acceleration (3/3)

We know that for the velocity there is the formula

$$v = |\vec{v}| = |\vec{\omega}|r = \omega r \quad (\text{this was derived in Appendix 2})$$

Putting this into Eq. (1) we get:

$$a = \frac{v^2}{r} = \frac{(\omega r)^2}{r} = \omega^2 r$$

$$\boxed{|\vec{a}| = |\vec{a}_n| = |\vec{\omega}|^2 r}$$

Centrifugal acceleration originates due to the non-inertial system and act with the same size, but in an opposite direction.